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Supplementary Material for

Earthquake in a Maze: Compressional Rupture Branching During the 2012 M_w 8.6 Sumatra Earthquake

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Published 19 July 2012 on *Science Express*

DOI: 10.1126/science.1224030

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Materials and Methods

Back-projection data selection and processing:

We processed the seismic data recorded by various seismic arrays at epicentral distances between 30 and 90 degrees. The large aperture and dense spacing of the European (www.orfeus-eu.org; Figure S1) and Japanese Hi-net networks (www.hinet.bosai.go.jp; Figure S2) provided fine spatial resolution for this event. The two arrays are located at almost orthogonal azimuths relative to the source area, providing complementary views of the rupture process. The P waveforms were filtered between 0.5 and 1 Hz, selected by their signal-to-noise ratio and by the mutual coherency of their initial 10 seconds, and then aligned by multi-channel cross-correlation. We applied the MUSIC back-projection technique (4, 31) on sliding windows ten seconds in length. This source imaging technique allows tracking the migration of the multiple sources of high-frequency (HF) radiation of an earthquake rupture. Our particular approach combines a high-resolution array processing technique (32, 33) with multi-taper cross-spectral estimation (34) to achieve higher resolution than conventional beamforming. Adopting a "reference window" strategy avoids the systematic "swimming" artifact (5). The MUSIC pseudo-spectrum is back-projected into the off-Sumatra region based on P travel times computed by the Tau-P toolkit and the IASP91 model (36)

Multiple point source analysis:

Seismic waves from the 2012 Sumatra earthquake sequence were recorded by a large number of broadband seismic stations enabling robust and reliable characterization of the overall source characteristics at long periods. For the $M_w=8.6$ event, 113 channels were used to perform a multiple point source inversion. We inverted simultaneously for the moment tensors of the subevents, their locations (latitude, longitude, depth) and time delays using a global sampling approach. The inversion is performed using the W-phase waveforms at shorter period (150-500 s) than the usual practice for single point source inversions of $M_w>8.0$ earthquakes. The optimal centroid depth of the mainshock is 30 km. Our preferred model is a two point source solution which is presented on Fig. 1.

Resolution and uncertainty of back-projection source imaging:

We conducted several analyses to understand the strengths and limitations of our back-projection method, as specifically applied to the European and Japanese arrays. We address two different array performance criteria. The spatial *resolution* of an array is defined as the capability to separate two simultaneous sources of different location: the *resolution length* is the minimum distance between sources that can be distinguished without ambiguity. The spatial *uncertainty* of an array is defined as the error in estimates of source location for isolated sources.

The resolution length that can be achieved with linear beamforming is conventionally estimated as the width at half peak amplitude of the main lobe of the array response function (ARF, 35). Figs. S4 and S5 show the ARFs of the Japanese and the European networks back-projected into the off-Sumatra region. The European ARF is relatively

compact with resolution lengths of 35 km in the N-S direction and 65 km in the E-W direction. These are one order of magnitude smaller than the source dimension of the M8.6 off-Sumatra event, which indicates that the overall rupture process can be resolved as long as there is adequate signal coherence. The Japanese Hi-Net array has a larger and more anisotropic ARF, with resolution lengths of 450 km in the N-S direction and 70 km in the E-W direction. This helps explaining why the early separation of the bilateral rupture fronts on fault B is imaged by the European array (movie S2) but the two fronts appear simultaneously in the Hi-Net images only when they reach the ends of fault B.

In back-projection source imaging with linear beamforming, the estimation of source location is based on identifying the peak position of an image that is smeared by the array response. The Cramer-Rao bound, a theoretical estimate of the upper bound of uncertainty of an estimator, indicates that the source location uncertainty is proportional to the resolution length of the ARF and inversely proportional to the signal-to-noise ratio of the images (SNR), defined as the ratio between the main peak amplitude and the mean amplitude of the background peaks of the image. The SNR depends on both the signal quality (noise and multipath) and the station density (sidelobe amplitude). The Hi-Net array has a particularly large SNR due to its small ambient noise level (borehole stations) and extremely dense spacing (~20 km).

However, due to the nonlinearity of the MUSIC technique, its resolution cannot be fully appreciated by considering the ARF. Typically, the resolution length of MUSIC is at least two times smaller than the beamforming resolution length (31). Moreover, the resolution analysis based on the ARF assumes perfect signal coherence. In practice, we quantify the uncertainty of back-projection of Hi-Net data by bootstrapping the back-projection of a M6.1 aftershock that occurred on April 21th, 2012 at 93.39 degree E and 2.22 degree N. The aftershock seismograms are first aligned by the initial 3 s of the P-wave arrival. The noise at each station is computed as the waveform residual with respect to the array-stacked waveform. The noise is shuffled by randomizing its Fourier phase spectrum, then added back to the stack. One thousand synthetic realizations of the aftershock array recordings plus noise are then back-projected considering the 10-s long window that begins at the initial arrival. The resulting back-projection locations are shown in Fig. S6. The bootstrap 95% confidence ellipse is elongated along the N-S direction with a major axis length of 9.5 km and minor axis length of 2.8 km. Unfortunately, the recording of this aftershock at the European array is too weak to be used in the 0.5-1 Hz frequency band.

To further understand the performance of the back-projection on complicated ruptures, we conduct two synthetic earthquake scenarios. To include key characteristics of the real wavefield, such as the decay of waveform coherence as a function of time, we use the M6.1 event as empirical Green's function. Inspired by our final back-projection result, particularly a jump from the fault C to fault D imaged with the Hi-Net array, we test scenarios of bilateral rupture on Fault A and B and unilateral rupture on Fault C with (Fig.S7) or without (Fig.S8) an additional fault D. We consider a uniform distribution of sub-sources, regularly located every 15 km along each fault. The rupture times correspond to an assumed rupture speed of 2.5 km/s. The MUSIC back-projection

technique recovers the location and timing of the scenario sources very well. The uncertainty of the peak locations is less than 20 km, which is reasonably good considering the coda and interference between sub-sources. These synthetic tests indicate that the jump between faults C and D is resolvable by the Hi-Net array.

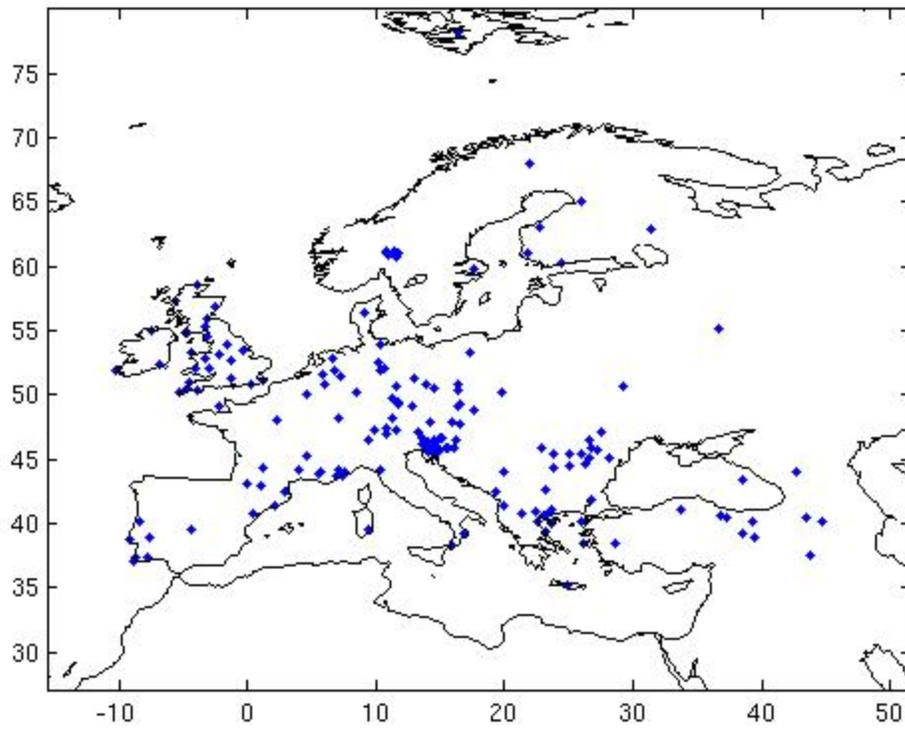


Fig. S1

Selected stations used for back-projection from the European network.

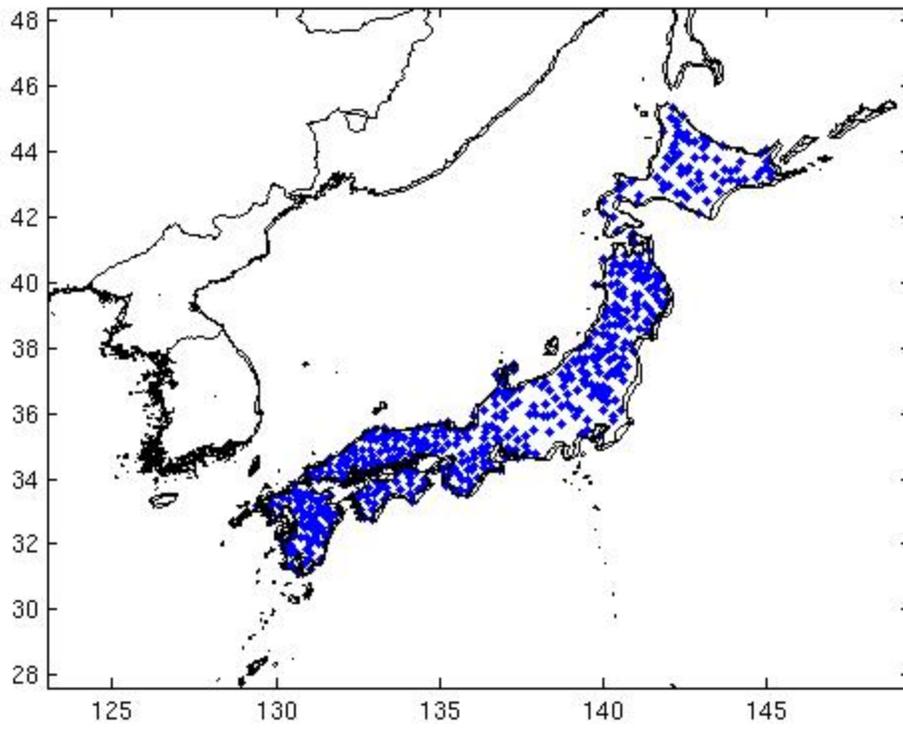


Fig. S2

Selected stations for back-projection from the Hi-Net (Japan) network.

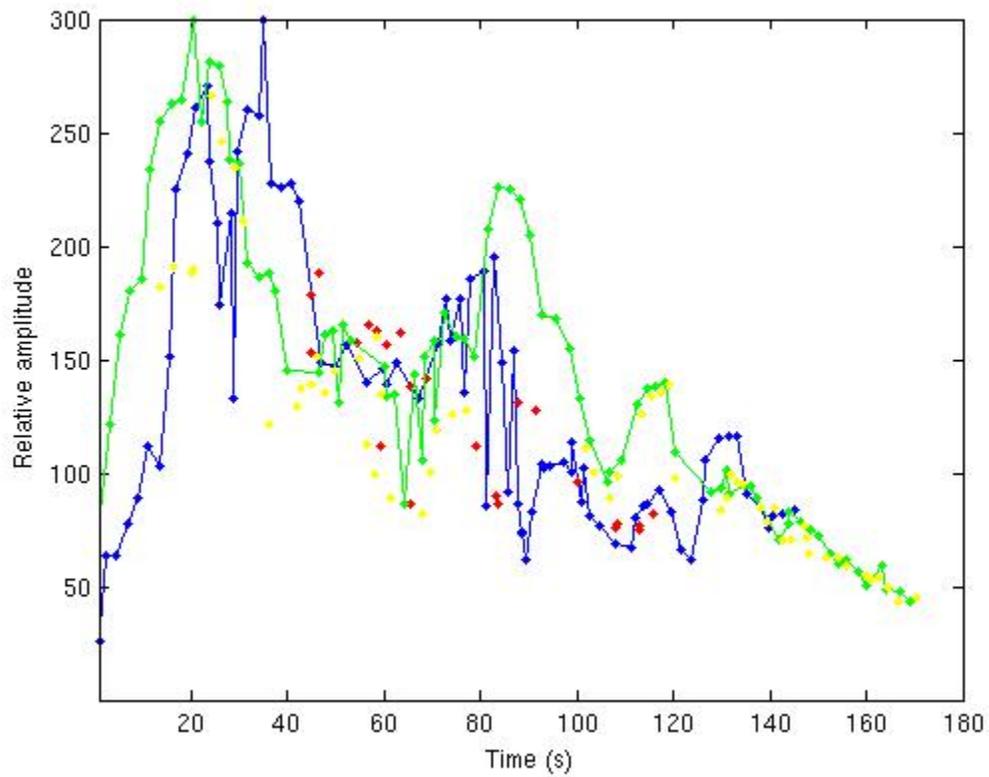


Fig. S3

Beamforming source amplitude evaluated at the location of the HF radiators obtained by MUSIC as a function of time, obtained with the European array (green) and Japanese Hi-Net array (blue). The scale is normalized by the beamforming amplitude of the initial window. The yellow and red dots indicate the amplitude of the secondary sources for the European and Japanese arrays, respectively.

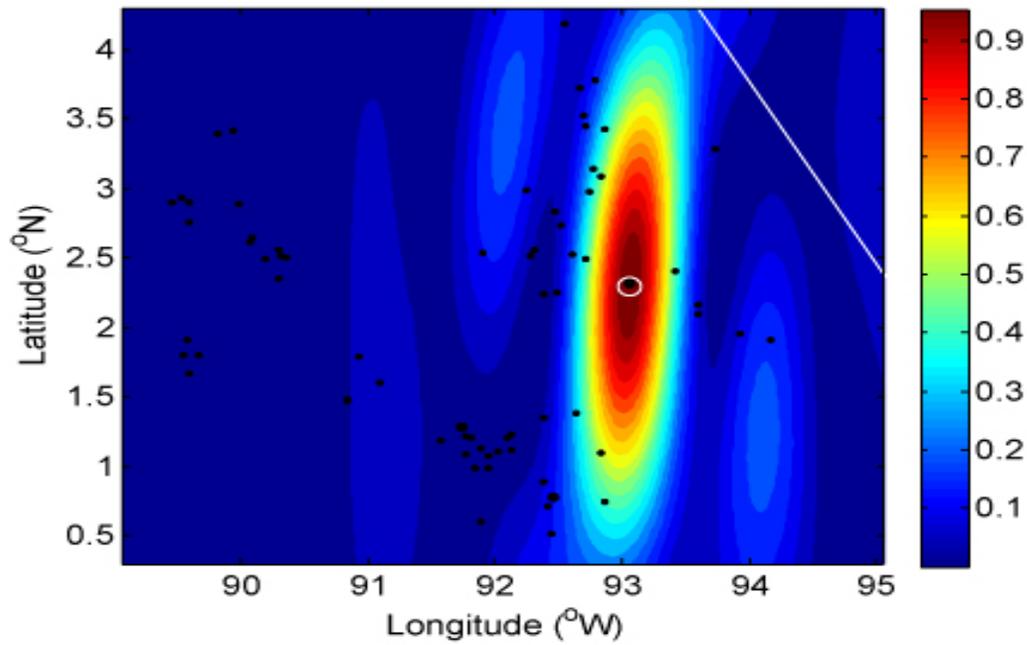


Fig. S4

Array response pattern of the Japanese Hi-Net array back-projected into the off-Sumatra earthquake region, plotted in map view. The color scale indicates the power of the array response, normalized by its peak value. The white line denotes the trace of the Sumatra trench. The white circle is the location of the epicenter. The black dots are early aftershocks that occurred within 24 hours of the mainshock.

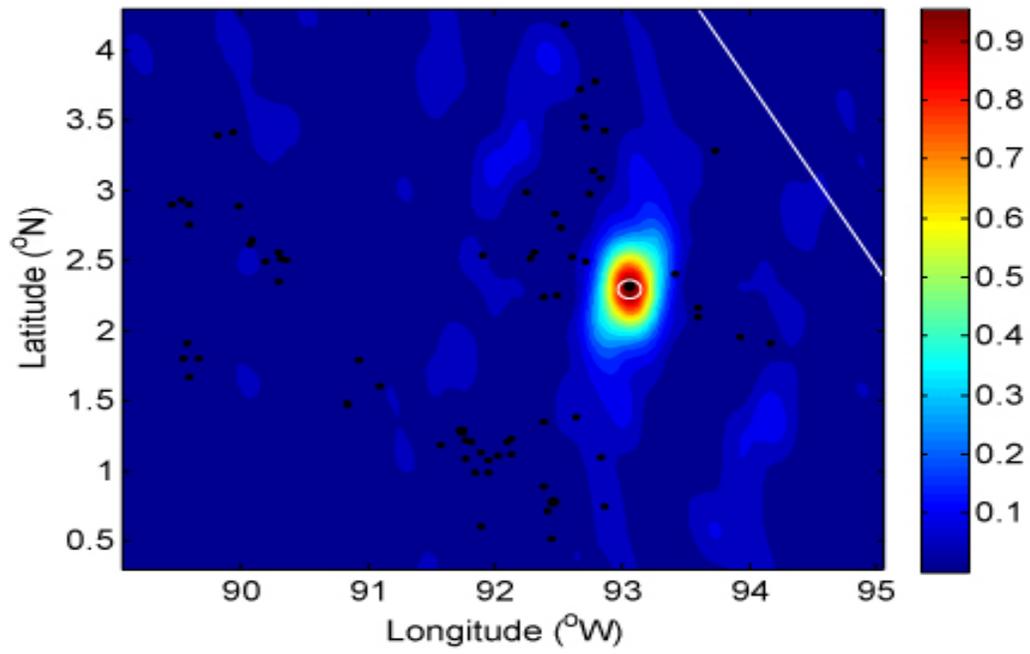


Fig. S5

Array response pattern of the European network. The convention is the same as in the previous figure.

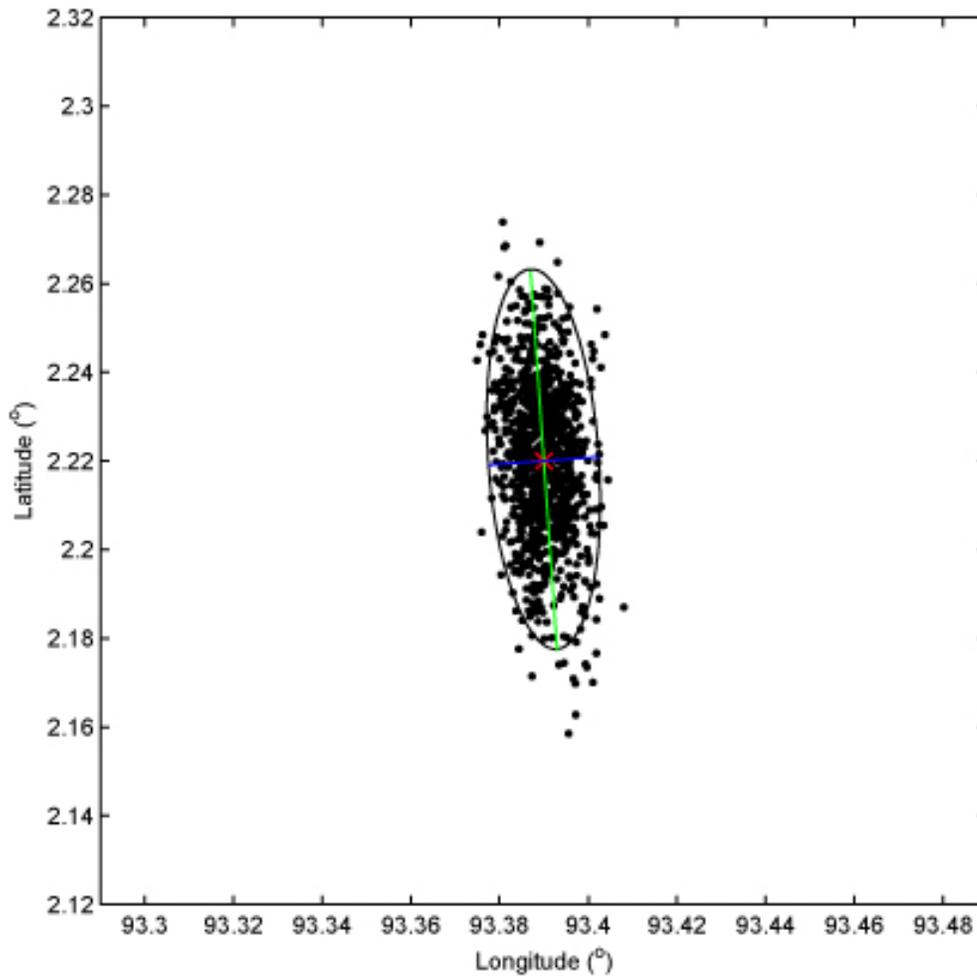


Fig. S6

Uncertainty of the back-projection. A M6.1 aftershock that occurred on April 21th, 2012 is used here as an empirical Green's function. The red asterisk denotes the hypocenter. The black dots are the back-projection of the bootstrapped aftershock recordings. The ellipse is the 95% confidence interval. It has a 9.5 km long major axis (blue line) and 2.8 km long minor axis (green line).

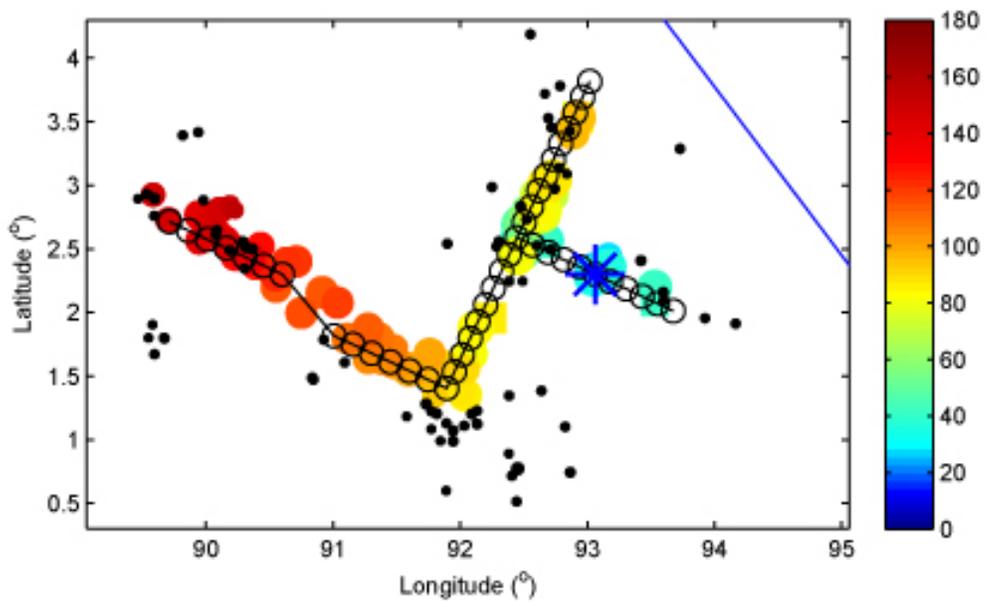


Fig. S7

Back-projection of a synthetic rupture scenario. The black circles are the synthetic sources. The colored circles are the recovered back-projection locations color-coded by time. The black dots are the early aftershocks that occurred within 24 hours of the mainshock.

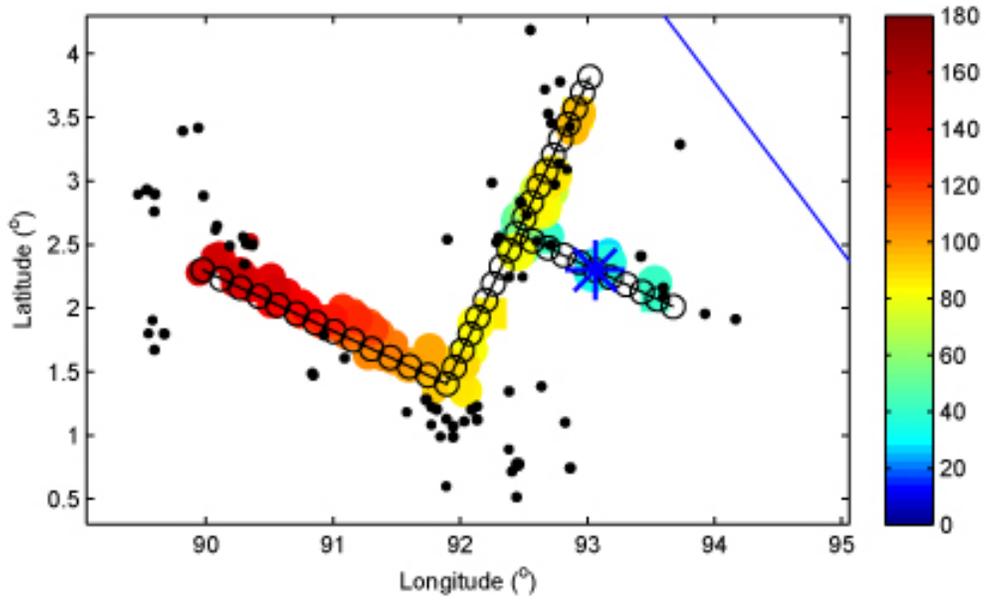


Fig. S8

Back-projection of a synthetic rupture scenario without offset between faults C and D. The convention of the figure is the same as in the previous one.

Movie S1

The movie shows the raw results of back-projection source imaging based on teleseismic data from European networks. Warm colors indicate the positions of the high frequency (0.5 to 1 Hz) radiation back-projected onto the source region based on IASP91 travel times (36). The sliding window is 10-s long and the origin time is 08:38:37, 04-12-12 (UTC). The beginning of the sliding window is set to be 5 s before the initial P-wave arrival. Colors indicate the amplitude of the MUSIC pseudo-spectrum on a logarithmic scale (dB) after subtracting the background level and rescaling the maximum to the linear beamforming power in each frame separately. The white star is the mainshock epicenter and the green circles are the epicenters of the first day of aftershocks from the NEIC catalog. Time relative to hypocentral arrival time is shown on top. The trench and coastlines are shown by white curves.

Movie S2

Back-projection source imaging based on teleseismic data from Japanese networks. Same convention as the previous movie.

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