Symmetrically reduced Galileon equations and solutions

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I. INTRODUCTION

Galileons, while popular in the cosmological literature, deserve to be better understood on their own. The present contribution is the study of their “maximal” forms in arbitrary dimension $D$, with a view to finding some interesting simplifications and explicit solutions of their (rather exotic) field equations. Such searches are usually most successful for systems with symmetry, as turns out to be especially true here. We will find dramatic reduction of both the effective equations and their solutions as the degree of symmetry is increased; thus, spherically symmetric configurations obey the $D=2$ Monge-Ampere equation rather than the $D$-dimensional Hessian one. This enables us to display quite general classes of explicit solutions that are independent of the models’ dimensionality, in contrast to the known general but quite implicit solutions.

We begin with a summary of the models; in the formulation of [1], the actions for the scalar field $\phi$ are given by

$$I = \int d^Dx \phi \left( e^{\mu\nu} \cdots e^{\gamma\delta} \cdots \phi_{,\mu\gamma} \phi_{,\nu\delta} \cdots \phi_{,\alpha\beta} \right)$$

$$= \int d^Dx \phi \det(\phi_{,\alpha\beta}).$$

(1)

The field equation states that the $D$-dimensional determinant of its second derivatives—i.e., the Hessian [2]—vanishes:

$$\det(\phi_{,\alpha\beta}) = 0.$$  

(2)

As already noted in [2], all $\phi$ that are independent of at least one Cartesian space-time coordinate are trivially solutions: the vanishing of any derivative annihilates a complete row and column of the determinant. This implies a sort of anti-Birkhoff theorem: since all static functions are solutions, interesting ones must depend on $t$. It also demonstrates the fact that one can have too much symmetry, here that of a cyclic coordinate. (Note also that, unlike for normal Lorentz invariant systems, rescaling each coordinate separately by an arbitrary constant is allowed because each term in the determinant expansion contains all derivatives.) We now turn to the reduction of (2) effected by less drastic symmetries.

II. SPHERICAL SYMMETRY

We begin with maximal, spherical symmetry, $\phi = \phi(r,t)$, and show that $\phi$ obeys the ($D=2$) Monge-Ampere equation for all $D$. (Missing angular dependence is not the same as missing Cartesian coordinates—in the former case, all the $x^i$ are still present, just in a particularly symmetric combination.) Note first that

$$\phi_{,i} = x_i/r \phi_{,i},$$

$$\phi_{,ij} = \delta_{ij}(\phi_{,r}/r) + x_i x_j/r^3(\phi_{,rr}/r - \phi_r),$$

(3)

where prime $= d/dr$, dot $= d/dt$. We write $\phi_{,\mu\nu}$ as the matrix

$$\phi_{,\mu\nu} = \left( \begin{array}{cc} \phi_{,i}/r & \phi_{,r}/r \\ \phi_{,r}/r & x_i x_j/r^3(\phi_{,rr}/r - \phi_r) \end{array} \right).$$

(4)

This implies that the full Hessian matrix (4) is of the form

$$\hat{S} = \left( \begin{array}{cc} \phi & b^T \\ b & 0 \end{array} \right).$$

(5)

with determinant

$$\det(\hat{S}) = \phi \det(\left[ D - b \phi^{-1} b^T \right]).$$

(6)

which is the standard result for block matrices. Using $b_i = \phi^i/\phi_r$ and $D_{ij} = \delta_{ij} x_j + x_i x_j \phi_{,rr}/r^3 - x_i x_j \phi_{,r}/r^2$, the matrix in (6) is

$$[D - b \phi^{-1} b^T]_{ij} = \delta_{ij} \phi_{,r}/r + x_i x_j \left( \phi_{,rr}/r - \phi_{,r}/r^3 \right).$$

(7)

But Sylvester’s theorem states that
\[
\det[\alpha \delta_{ij} + \beta x_i x_j] = \alpha^{D-1} \left( 1 + \frac{\beta}{\alpha} r^2 \right),
\]
\[
\det[D - \mathbf{b} \hat{\phi}^{-1} \mathbf{b}^T] = \left[ \frac{\phi''}{r} \right]^{D-1} \left( 1 + \frac{r}{\phi'} (\phi'' - \frac{\phi'''}{r} - \frac{\phi''}{\phi} \phi') \right)
\]
\[
= \left[ \frac{\phi''}{r} \right]^{D-2} \left( \phi'' - \frac{\phi'''}{\phi} \phi' \right). \tag{9}
\]
By (6), and dropping the overall coefficient, the reduced field equation is
\[
(\phi'' \hat{\phi} - \phi''') = 0, \tag{10}
\]
which is independent of \( D \), as promised. In the most degenerate, one-variable case \( \hat{\phi} = \phi[\sqrt{x^2 + \sqrt{t^2 + r^2}}] \), the equation reduces to \( \phi'' = 0 \), whose solution is just \( \phi = a\sqrt{x^2 + b} \); this is of course distinct from the “pure gauge” \( \phi = a_\mu x^\mu + b \), whose second derivatives vanish identically. This solution exemplifies the result [3] that any function homogeneous of order one in \( (t, r) \) is a solution.

It is not difficult to guess some classes of interesting solutions of (10). In particular, there are traveling solutions, \( \phi(t, r) = \Phi(at + br) \). Indeed, superpositions of traveling solutions, \( \phi = \Phi_1(at + br) + \Phi_2(ct + dr) \), seem to exist, subject “only” to the condition \( ad = bc \); unfortunately, this condition reveals the superposition to be the initial \( \phi = \Phi(at + br) \) in disguise. Recalling the freedom of rescaling \( (t, r) \) by arbitrary constants, we choose to remove \( a \) and \( b \) from \( \Phi_1 \), which then makes \( \Phi_2 = \Phi_2(c/at + d/br) \); however, since \( ad = bc \), \( \Phi_2 \) is again a function of just \( t + r \), like the (rescaled) \( \Phi_1 \) and \( \phi = \Phi(t + r) \). Our traveling wave broadens, in a sense, the homogeneity theorem of [3]: obviously, any—even independent—rescalings of \( t \) and \( r \) are permitted for \( \Phi(at + br) \), as indeed is the case for their lower symmetry extensions studied below.

Actually, the general, if highly implicit, solution of (10) is known and can be correlated to some of the explicit solutions: we follow [3,4]. The idea is to generate \( \phi(t, r) \) whose \( \phi' \) and \( \phi'' \) are related; then, the columns of the Hessian matrix will not be independent, and hence the determinant of the matrix will vanish. One can accomplish this by setting \( \phi(t, r) = f(u), \phi'(t, r) = g(u) \), where \( f \) and \( g \) are some functions of the single variable \( u(t, r) \). For then
\[
\begin{align*}
\phi & = f'(u) \dot{u}, \\
\phi' & = f''(u) u', \tag{11}
\end{align*}
\]
and the combination in (10) vanishes. In order to generate a \( \phi \) with the appropriate derivative relations, we take
\[
\phi(t, r) = tf(u) + rg(u) + ua, \tag{12}
\]
for an arbitrary constant \( a \). Then
\[
\dot{\phi} = f' + u(t' + r') + a, \\
\phi' = g + u'(t' + r' + a),
\]
so that \( \phi = f \) and \( \phi' = g \) hold if
\[
t(t' + r') + a = 0. \tag{14}
\]

The solutions (12) and (14) must both be satisfied, and together they implicitly define the most general solution to (10). The freedom in the choice of \( f \) and \( g \) can be used to reproduce explicit solutions. Perhaps the easiest of these is the single-variable one found earlier: taking \( (f, g) = (\cos \mu, \sin \mu) = (t, r)/\sqrt{t^2 + r^2} \) yields the desired \( \phi = \sqrt{t^2 + r^2} \); hyperbolic functions yield \( \phi = \sqrt{t^2 - r^2} \).

\section*{III. LOWER SYMMETRIES}

It is rather clear what to expect as the symmetry level decreases. Consider the axial symmetry, which depends on one axial direction (the \( z \) direction) and the radial variable \( \rho^2 = x_1^2 + \cdots + x_P^2 \); so, \( \phi = \phi(t, \rho, z) = \phi(Y_1) \). We omit the details, but from the spherical discussion, one expects the \( D = 3 \) determinant extension of Monge-Ampere,
\[
\det \left( \frac{\partial^2 \phi}{\partial Y_i \partial Y_j} \right) = 0. \tag{15}
\]
This is indeed borne out by Sylvester-like decomposition of the full \( \phi_{\mu \nu} \) into a \( D = 2 \) \(( t, z)\) times the “planar” spherically symmetric \( D - 2 \).

The solutions are again obtainable in a general but implicit form. It is perhaps not surprising that \( \phi = a\sqrt{t^2 + \rho^2 + z^2} + b \) and \( \phi = \Phi(at + b\rho + cz) \) are also solutions. In surprising contrast to the spherical result, however, we find here that superposition, \( \phi = \Phi_1(at + b\rho + cz) + \Phi_2(dt + ep + fz) \), holds: unlike the spherical case, there is no constraint among the constants.

Finally, we derive the effective reduced Hessian equations for arbitrary symmetry configurations by writing (2) in terms of the appropriate generalized coordinates. Let us call, collectively, the respective relevant and missing directions “\( r \)” and “\( \theta \)”, with the boldfaces indicating that there can be more than one. The number \( R \) of \( r \) directions determines the dimension of the reduced \( (t, r) \) Hessian. Generalized coordinates of course require the introduction of covariant derivatives, replacing \( \partial_i \partial_j \) by \( D_i D_j \) on our scalar. (The \( D_i \) commute in flat space, and there is no change in the time derivatives.) The Hessian equation then becomes
\[
\epsilon^{i\alpha \theta} \epsilon^{j\rho \phi} D_i \phi \cdots D_j \phi = 0. \tag{16}
\]
Its dimension is still \( D \), but it effectively contracts after we implement the symmetries: the factors \( D_i \phi \) reduce to \( \partial_i \phi \). The \( D_r \phi_\theta \) or \( D_\theta \phi_r \sim \Gamma^{\alpha \theta \phi}_\theta \phi_r \phi_\theta \) vanish: the only nonzero connections have two \( \theta \) and one \( r \), while \( \partial_\theta \phi = 0 \); also, \( D_r \partial_r = \partial_r \partial_r \). There remains \( D_{\theta \phi} \partial_r \phi \sim \Gamma^{\theta \phi_\theta \phi}_\theta \phi_r \phi_\theta \partial_r \phi \).
multiplied by \( r \) and an angular coefficient. Further, \( D_\theta \partial_\theta \phi \) is a common factor to every term of (16); it provides the (only) nonvanishing contraction of the \( \theta \) indices in \( \epsilon \epsilon \). Precisely because it is common to each term in (16), it may be canceled out of the equation. So (16) reduces, as predicted, to a determinant of dimension \( R + 1 \) (there are of course \( D - 2 - R \theta \)), which is formally like our spherical result (10), but now the vector \( r \) determines the number of \( \phi \) factors there: quadratic for spherical, cubic for axial symmetries, etc. Specifically, in our vector notation, the final equation is just

\[
[\ddot{\phi} \partial_r \partial_r \phi - \dot{\phi} \partial_r \dot{\phi} ] = 0,
\]

whose (\( R + 1 \)) dimension is implicit in our vector notation.

**IV. SUMMARY**

We have been able to tame the general \( D \), maximal Galileons for configurations with arbitrary degrees of symmetry, thereby gaining some concrete view of their explicit behavior beyond the (rather forbidding) implicit general solutions of the full Hessian equation. Similar reductions could be envisaged in terms of, e.g., light cone coordinate choices where \( t \) becomes mixed with spatial variables. Ours can hardly pretend to be a detailed analysis, however. We have not attempted to seek solutions with proper asymptotic behavior, let alone study the necessarily nonminimally coupled Galileon-gravity equations—a qualitatively harder challenge, if perhaps still feasible in the spherical limit.

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