# Magnetic Hyperfine Structure and Core Polarization in the Excited States of Lithium\*

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The magnetic hyperfine splitting constants,  $a_J$ , from theoretical calculations on the  $3^2S$ ,  $2^2P$ ,  $3^2P$ , and  $3^2D$  excited states of the Li atom are reported. The wave functions were calculated using the author's GF method, which corresponds to optimizing the orbitals of a Slater determinant after spin projection. Thus the wave functions include core polarization, but no appreciable correlation. For the  $2^2P$  state we calculate  $a_{1/2} = 0.2206$  a.u.  $(45.74 \text{ Mc/sec for }^7\text{Li})$  which is in good agreement with the experimental value,  $0.2227 \pm 0.0017$  a.u.  $(46.17 \pm 0.35 \text{ Mc/sec for }^7\text{Li})$ , and the value from configuration interaction calculations, 0.2206. Thus for the Li atom core polarization accounts for most of the error in the Hartree–Fock values of  $a_J$ . These calculations yield  $\langle 1/r^3 \rangle$  and spin density, Q(0), in agreement with other accurate theoretical calculations, and in disagreement with the values found using level crossing experiments, indicating that the interpretation of the level crossing experiments in terms of  $\langle 1/r^3 \rangle$  and Q(0) may not be correct. It is found that for both the unrestricted Hartree–Fock and GF methods the use of a different Hamiltonian for each electron leads to virtual orbitals which are good approximations to the actual orbitals of excited states.

#### INTRODUCTION

The Fermi contact portion of the magnetic hyperfine interaction is proportional to the spin density at the nucleus, Q(0). The Hartree-Fock (HF) wave function for an atom such as Li  ${}^{2}P(1s^{2}2p)$ , N  ${}^{4}S(1s^{2}2s^{2}2p^{3})$ , or O  ${}^{3}P(1s^{2}2s^{2}2p^{4})$  has all s orbitals doubly occupied and hence leads to a zero Fermi contact interaction. However, the experimental results show that the Fermi contact term is not zero in these atoms<sup>1,2</sup> and thus that the spin density at the nucleus is nonzero. This clearly indicates that to study such interactions we must improve upon the Hartree-Fock wave function. One approach has been to use configuration interaction (CI), 3,4 where many Slater determinants are formed from a large set of orbitals, and the determinants are combined such as to minimize the total energy. If this is carried out for a complete set of orbitals, we must get the exact wave function, and thus for a sufficiently large basis set we should be able to get arbitrarily good energies and spin densities and correctly account for the magnetic hyperfine interactions.

Another approach has been to retain the Slater determinant form of the wave function, but to remove some of the other restrictions of the Hartree-Fock method. In particular if we allow the 1s orbitals of the Li 2P wave function to be different for different spins, there will be a net spin density at the nucleus,  $|\phi_{1a}(0)|^2 - |\phi_{1b}(0)|^2$ , where  $\phi_{1a}$  is associated with  $\alpha$  spin and  $\phi_{1b}$  with  $\beta$  spin. This procedure, which is called the unrestricted Hartree-Fock (UHF) method, 5 has the disadvantage6 that the many electron wave function is not an eigenfunction of  $\hat{S}^2$ . Since the property we are considering depends on spin, we are led to consider the spin projected UHF (PUHF) wave function in which we delete the parts of the UHF wave function having the wrong spin. However, in this case we are using the orbitals optimized for the unprojected (UHF) function, whereas we really should use the orbitals optimized for the projected function. This latter approach in which the orbitals are optimized after spin projection is equivalent to the GF method<sup>7-9</sup> for electronic wave functions. That is the GF wave function is the best possible spin projected Slater determinant wave function.<sup>7</sup>

Since the GF wave function is based on spin projection of a single Slater determinant, it will not coincide with the exact wave function and will not yield the spin density exactly. However, it does take into account the direct spin polarization of the core (we will call this core polarization) caused by the unpaired spins of the valence electrons. Any difference between the GF value of Q(0) and the exact value could be accounted for by mixing in sufficient other spin projected determinants; such effects we will simply ascribe to correlation. Thus we distinguish the simple exchange correlation involved in core polarization and accounted for by GF wave functions from the other kinds of correlation.

We will see that just as for the Li  $2^2S$  ground state<sup>8</sup> the GF wave function also accounts for the hyperfine structure of the Li  $2^2P$  state, and thus corrections beyond direct core polarization are not too important for Li.

### **CALCULATIONS**

The  $2^2S$ ,  $3^2S$ ,  $2^2P$ ,  $3^2P$ , and  $3^2D$  states of Li were considered, and UHF, PUHF, and GF calcutions were carried out for each state. Since HF wave functions for these states have been published by Weiss, <sup>4</sup> we used basis sets based upon his calculations. These basis sets of Slater orbitals <sup>10</sup> were selected so as to satisfy the cusp condition at the nucleus [this seems to be important for reliable values of Q(0), especially for larger atoms such as N]. It was established that for Li the optimum orbital exponents for all three types of wave

functions (HF, UHF, and GF) are essentially the same. Four of the ns orbitals were primarily used by the core electrons and had been taken to have the same orbital exponents in all of Weiss's calculations. We reoptimized these orbitals for the 2 <sup>2</sup>P state and found that the new orbital exponents also led to lower energies and better virial ratios for the other states. Thus we changed Weiss's basis sets by using these reoptimized (for the 2  $^2P$ state) exponents for the core orbitals.

A self-consistent-field calculation (e.g., HF, UHF, or GF) for the second-lowest state of some symmetry [e.g.,  $3^{2}P(1s^{2}3p)$ ] is straightforward and leads to an upper bound on the exact energy of the excited state (see Appendix C). However, for the 32S and 32P states of Li the solutions of the

usual form of the UHF and GF equations diverged and oscillated. This required using the form of the self-consistent-field equations derived directly from the variational theorem rather than the usual form of the equations in which self-repulsion terms are added. This is discussed later.

### CORE POLARIZATION

The calculated hyperfine splitting constants are given in Table I (see Appendix B for the relationships between  $a_J$  and  $\hat{Q}(0)$  and  $\langle 1/r^3 \rangle$ ). As occurred for the 2  $^2$ S state,  $^7$  the  $a_{1/2}$  for  $2^2P$  as calculated by both the GF and CI<sup>11</sup> methods is in good agreement with experiment. 12 Since the GF results are close to experiment, we see that the main error in the

TABLE I. Hyperfine splitting constants and other properties for the ground and excited states of Li as obtained from GF and other wave functions. E is the total energy, V is the potential energy, V/2E is the virial ratio,  $\rho(0)$  is the electron density at the nucleus, and Q(0) is the spin density at the nucleus. All quantities are in atomic units (the hartree is the unit of energy).

	E	V/2E	$\rho(0)$	$\langle \sum r^2 \rangle$	Q(0)	$\langle 1/r^3 \rangle$	$a_{L-\frac{1}{2}}$	$a_{L+\frac{1}{2}}$
$2^2S$ HF	<b>-</b> 7.432725	1.000 002	13.8160	18.638	0.1667			1.396
UHF	-7.432749	1.000 000	13.8159	18.630	0.2248		• • •	1.883
PUHF	-7.432767	0.999999	13.8159	18.630	0.1866		• • •	1.563
$\mathbf{GF}$	<b>-7.432813</b>	1.000000	13.8159	18.609	0.2406			2.016
CI	-7.477 9 <sup>a</sup>				$0.2249^{b}$			1.884
EXP	-7.478 07 <sup>C</sup>				0.2313 <sup>d</sup>	•••	•••	1.937
3 <sup>2</sup> S HF	-7.310 210	0.999994	13.7080	119.480	0.038 64		• • • •	0.3237
UHF	-7.310 212	1.000002	13.7067	119.470	0.05253	•••	•••	0.4401
PUHF	-7.310 213	1.000000	13.7067	119.469	0.04335	• • •		0.3632
GF EXP	-7.310216 $-7.35410$	1.000 002	13.7067	119.447	0.056 22	• • •	•••	0.4710
$2^2P$ HF	-7.365 069	0.999995	13,6534	28.716	0.00000	0.05848	0.1559	+0.0311
UHF	-7.365 076	1.000 000	13.6535	28.709	-0.01747	0.05852	0.2407	-0.0176
PUHF	-7.365 080	1.000 000	13.6535	28.709	-0.00582	0.05852	0.1722	+0.0149
$\mathbf{GF}$	-7.365 091	1.000002	13.6534	28.692	-0.023 04	0.05861	0.2206	-0.0331
CI	-7.40838 <sup>e</sup>				-0.02222	g	0.2206	-0.0302
EXP	$-7.41016^{f c}$						0.2227 <sup>h</sup>	i
$3^2P$ HF	-7.293 186	0.999945	13.6660	172.536	0.00000	0.01759	0.04689	+0.00937
UHF	-7.293 187	0.999972	13.6661	172.530	-0.005531	0.01760	0.06234	-0.006 07
PUHF	-7.293 188	0.999972	13.6661	172.530	-0.001843	0.01760	0.05204	+0.004 22
$_{\rm EXP}^{\rm GF}$	-7.293 189 -7.337 15 <sup>C</sup>	0.999972	13.6661	172.511	-0.007318	0.01759	0.067 35	-0.01106
$3^2D$ HF	-7.291 973	0.999936	13.6715	126.924	0.00000	0.004 951	0.004 948	0.001412
UHF	-7.291 974	0.999966	13.6722	126.923	-0.000036	0.004951	0.005 008	0.001352
PUHF	<b>-</b> 7.291 974	0.999966	13.6722	126.923	-0.000012	0.004 951	0.004 968	0.001392
GF EXP	-7.291 974 -7.335 52 <sup>c</sup>	0.999966	13.6722	126.923	-0.000053	0.004 951	0.005 037	0.001 323

<sup>&</sup>lt;sup>a</sup>E. A. Burke, Phys. Rev. <u>130</u>, 1871 (1963). <sup>b</sup>E. A. Burke, Phys. Rev. <u>135</u>, A621 (1964).

<sup>&</sup>lt;sup>c</sup>Atomic Energy Levels, compiled by C. E. Moore, National Bureau of Standards Circular No. 467 (U.S.

Government Printing Office, Washington; D. C., 1947).

dP. Kusch and H. Taub, Phys. Rev. <u>75</u>, 1477 (1949); R. G. Schlect and D. W. McColm, Phys. Rev. 142, 11

<sup>(1966).</sup> The conversion factor from Mc/sec to a.u. is given in Appendix B.

e<sub>Weiss</sub>, Ref. 4

fArdill and Stewart, Ref. 18.

 $<sup>\</sup>frac{g}{(1/r^3)_{\text{orb}}} = .05975$ ,  $(1/r^3)_{\text{dip}} = .05923$  Ref. 17.

hRitter, Ref. 15.

<sup>&</sup>lt;sup>i</sup>See discussion in the Core Polarization section.

HF calculations for the magnetic field at the nucleus,  $\vec{H}(0)$ , is just the lack of core polarization. The error due to correlation is the difference between the GF and experimental values which we see is small.

The  $a_{3/2}$  hfs constant for the 2  ${}^2P$  state could not be measured directly. 12 However, some level crossing experiments were carried out  $^{13}, ^{14}$  which led to effective values of  $\langle 1/r^3 \rangle$  and Q(0) that could be extrapolated to zero magnetic field in order to calculated  $a_{3/2}$ . The resulting value for  $\langle 1/r^3 \rangle$  is 0.0645 a.u. which differs from the CI value<sup>5</sup>, 15 by about 8%. The CI values of 0.0596 and 0.0594 for the  $a_{1/2}$  and  $a_{3/2}$  states are consistent with HF, UHF, PUHF, and GF values of 0.0585 to 0.0586 and should be expected to be within 1\% of the exact value. Since the CI and GF values of Q(0) and  $\langle 1/r^3 \rangle$  are consistent with the experimental  $a_{1/2}$ , it may be that the interpretation of the experimental results in terms of Q(0) and  $\langle 1/r^3 \rangle$  are incorrect. <sup>15</sup> In particular the external field, H. must lead to perturbations in the core orbitals which depend on H and thus lead to effective values of Q(0) and  $\langle 1/r^3 \rangle$ which are different for each level and which depend on H.  $^{14}$  Since the value of  $a_{3/2}$  from the GF and CI calculations is a factor of two larger than that extrapolated from level crossing experiments, it may be that direct measurements at small field could determine which of these values is best even though it cannot  $^{12}$  give a precise value to  $a_{3/2}$ . (Note from Table I that the HF and PUHF wave functions lead to the wrong sign of  $a_{3/2}$  for the Li  $2^{2}P$  and  $3^{2}P$ states.)

We should note here that since the UHF wave function is not an eigenstate of  $\hat{S}^2$  or  $\hat{J}^2$ , the UHF values for  $a_J$  are somewhat suspect since  $a_J$  is necessarily defined for some specific value of J. We have used the usual<sup>13,16</sup> (apparently reliable) procedure of pretending that the UHF wave function has the same J and S that it would have if the core orbitals were identical.

### Perturbation Approximations

Rather than calculate the effect of core polarization self-consistently as in the GF method, several perturbation approaches (starting with the HF wave function) have been suggested. 17,18 One interesting approach<sup>18</sup> lets the core orbitals be perturbed by the nuclear magnetic moment via the Fermi contact term. By solving for the electronic energy for the perturbed wave functions, one gets directly the core polarization contribution to  $a_{I}$ . Since the orginal perturbation is due to the nuclear moment, the distortion of the core orbitals should be relatively independent of the electronic state. Thus the core polarization in molecules or metals can be calculated more easily. This approach leads to an approximation to the UHF wave function and for the 2s and 2p states of Li leads to core polarization contributions to Q(0) of 0.0482 and -0.0154, respectively, which are about 85% of the UHF values. Similarly, a perturbation approximation 19 to the GF function for the 22S state yields a core polarization contribution to Q(0) of 81% of the GF value. Thus, self-consistency is important for a reliable value of Q(0).

#### **Summary on Core Polarization**

The term core polarization refers only to changes which occur upon improving the wave function within the independent particle interpretation, i.e., as in going from HF to UHF or HF to GF. The GF results given here show clearly the importance of core polarization and thus justify the concept both for polarization by s and non-selectrons. 20 Regardless of how the remaining correlation energy is taken into account, it leads to a minor change in Q(0). By comparing the GF and HF energies, we see that although the core polarization effects are crucially important for Q(0), they have an almost negligible (but numerically significant) effect on the energy, which is the reason that some very good (low-energy) calculations have yielded rather poor spin densities. 21

### CALCULATIONS ON EXCITED STATES

In carrying out calculations on the  $3\,^2S$  and  $3\,^2P$  excited states of Li, we observed some interesting features which would be expected to occur generally when performing calculations on excited states of the same symmetry as some lower state. In order to keep the discussion clear and simple, we will discuss the  $3\,^2S$  state of Li.

In the UHF method for Li we consider a wave function of the form

$$\alpha \left[ \phi_{1a}^{(1)} \phi_{2a}^{(2)} \phi_{1b}^{(3)} (3) \alpha(1) \alpha(2) \beta(3) \right],$$

and require that the orbitals be optimum, that is, yield a stationary energy. (In the above expression  $\alpha$  is the antisymmetrizer, and  $\alpha$  and  $\beta$  are the spin-up and spin-down eigenfunctions of  $\hat{s}_{Z}$ .) The result is a set of three equations

$$H_{1a}\phi_{1a} = \epsilon_{1a}\phi_{1a},$$

$$H_{2a}\phi_{2a} = \epsilon_{2a}\phi_{2a},$$

$$H_{1b}\phi_{1b} = \epsilon_{1b}\phi_{1b},$$

$$(1)$$

the self-consistent solutions of which are the optimum orbitals. In (1)

$$\begin{split} H_{1a} &= h + \mathfrak{I}_{2a} - \mathfrak{K}_{2a} + \mathfrak{I}_{1b}, \\ H_{2a} &= h + \mathfrak{I}_{1a} - \mathfrak{K}_{1a} + \mathfrak{I}_{1b}, \\ H_{1b} &= h + \mathfrak{I}_{1a} + \mathfrak{I}_{2a}, \end{split} \tag{2}$$

where

$$\begin{aligned} &\vartheta_{i}(1) = \int \! dx_{2} \phi_{i}^{*}(2) \phi_{i}(2) / r_{12} \\ & \varkappa_{i}(1) = \int \! dx_{2} \phi_{i}^{*}(2) \hat{P}_{12} \phi_{i}(2) / r_{12} \end{aligned}$$

are the Coulomb and exchange operators and  $\hat{P}_{\rm 12}$  permutes electrons 1 and 2. However,

$$(\mathfrak{I}_i - \mathfrak{K}_i)\phi_i = 0.$$

So (2) are usually replaced by

$$\begin{split} H_{A} &= h + (\mathfrak{I}_{1a} - \mathfrak{K}_{1a}) + (\mathfrak{I}_{2a} - \mathfrak{K}_{2a}) + \mathfrak{I}_{1b} \quad , \\ H_{B} &= h - \mathfrak{I}_{1a} - \mathfrak{I}_{2b} - (\mathfrak{I}_{1b} - \mathfrak{K}_{1b}), \end{split} \tag{3}$$

where (3) has the same self-consistent solutions as (2),

$$\begin{split} &H_{A}\phi_{1a}=\epsilon_{1a}\phi_{1a},\\ &H_{A}\phi_{2a}=\epsilon_{2a}\phi_{2a},\\ &H_{B}\phi_{1b}=\epsilon_{1b}\phi_{1b}. \end{split} \tag{4}$$

Besides the self-consistent solutions (i.e., the ones used in the field terms), (2) and (3) have an infinite number of other solutions, called the virtual orbitals. Let us consider the ground state of Li,  $2 \, ^2S[(1s)^2(2s)]$ . Then 1a and 1b correspond to 1s orbitals and 2a corresponds to the 2s orbital. From (1) and (2) we have

$$\epsilon_{2a} = \langle \phi_{2a} | h | \phi_{2a} \rangle$$

$$+ \mathfrak{g}_{1a,2a} - \mathfrak{K}_{1a,2a} + \mathfrak{g}_{1b,2a} , \qquad (5)$$

where

$$\begin{split} \mathcal{I}_{1a,2a} &= \langle \phi_{2a} | \mathcal{I}_{2a} | \phi_{1a} \rangle \\ &= \langle \phi_{2a}(1) \phi_{1a}(2) | 1/r_{12} | \phi_{2a}(1) \phi_{1a}(2) \rangle \end{split}$$

$$\begin{split} \mathfrak{K}_{1a,2a} = & \langle \phi_{2a} | \mathfrak{K}_{1a} | \phi_{2a} \rangle \\ = & \langle \phi_{2a}^{(1)} \phi_{1a}^{(2)} | 1/r_{12} | \phi_{1a}^{(1)} \phi_{2a}^{(2)} \rangle. \end{split}$$

Hence (5) is just the difference in energy between the three-electron wave function  $\alpha\left(\phi_{1a}\phi_{2a}\phi_{1b}\alpha\alpha\beta\right)$  and the two-electron wave function  $\alpha\left(\phi_{1a}\phi_{1b}\alpha\beta\right)$ , and thus (5) is expected to be close to the ionization energy for the 2s electron (Koopmans' theorem²²). Now consider the solution of  $H_{2a}$  which has two nodes and is spherically symmetric (i.e., the 3s virtual orbital), and call this orbital  $\phi_{3s}^{2a}$ 

$$H_{2a}\phi_{3s}^{2a} = \epsilon_{3s}^{2a}\phi_{3s}^{2a}.$$
 (6)

Then from (2) and (6)

$$\epsilon_{3s}^{2a} = \langle \phi_{3s} | h | \phi_{3s} \rangle + \mathfrak{g}_{1a,3s} - \mathfrak{K}_{1a,3s} + \mathfrak{g}_{1b,3s},$$
 (7)

which is just the difference between the energy of the three-electron wave function  $a(\phi_{1a}\phi_{3s}^{\phantom{3s}})^{2a}\phi_{1b}^{\phantom{3s}}\alpha\alpha\beta)$  and that of the two-electron wave function

 $\alpha(\phi_{1a}\phi_{1b}\alpha\beta)$ . Thus (7) is expected to close to the ionization energy for a 3s electron, and the virtual solutions of  $H_{2a}$  can be expected to be good approximations to the excited state orbitals of the Li atom.

Now consider the 3s virtual solution of  $H_A$ 

$$H_A\phi_{3s}^{\quad A}=\epsilon_{3s}^{\quad A}\phi_{3s}^{\quad A}\quad .$$

From (3) this has an energy of

$$\epsilon_{3s}^{A} = \langle \phi_{3s} | h | \phi_{3s} \rangle + \mathfrak{I}_{1a,3s} - \mathfrak{K}_{1a,3s}$$
  
+  $\mathfrak{I}_{1b,3s} + \mathfrak{I}_{2a,3s} - \mathfrak{K}_{2a,3s}$ .

But the ionization potential for the 3s electron should be as in (7); thus  $\epsilon_{3S}^A$  is too large by  $\mathfrak{I}_{2a,3s}-\mathfrak{K}_{2a,3s}$ , which is just the term arising from the addition of the self-Coulomb and exchange operators to  $H_{2a}$ . Since  $\mathfrak{I}_{2a,3s}>\mathfrak{K}_{2a,3s}$  the energies of the A-type virtual orbitals are all to high (in the case of the Li atom ground state, the virtual orbitals are all unbound), and the orbitals are too diffuse. This occurs because the virtual orbitals see an effective field due to N other electrons rather than N-1 other electrons. (The  $\phi_{3s}^A$  virtual orbital of  $H_A$  is an approximation to the  $\phi_{3s}$  in  $\mathfrak{a}(\phi_{1a}\phi_{2a}\phi_{3s}\phi_{1b}\alpha\alpha\alpha\beta)$ , which is a wave function for the 3  $^3S$  state of Li $^-$ , probably an unstable state.)

To illustrate these points, both (1) and (4) were solved for the ground state of Li using a basis set of 15 basis functions. The resulting eigenvalue spectra are given in Table II, where they are compared with the experimental and self-consistentfield values for the excited states. We see that the virtual solutions of  $H_{2a}$  are comparable to the self-consistent solutions for the excited states, whereas this is not at all true for the virtual solutions of  $H_A$ . The errors in the higher eigenvalues of  $H_{2a}$  are mainly due to the incompleteness of the basis set. If a complete basis set had been used, we probably would have obtained good values for the energies of all observed ns, np, nd, and nfstates of Li. Thus if one wishes to study a series of rydberg states of an atom or molecule without doing each state self-consistently, the HF or UHF equations analogous to (1) should be used, rather than those analogous to (4). Actually, the core orbitals can be solved for with (3) as long as the orbital to be excited is solved for with the  $H_i$  from

If we wish to solve self-consistently for the  $(1s)^2$   $(3s)^2S$  excited state of Li, we just solve (1) where the  $\phi_{2a}$  in  $\mathfrak{I}_{2a}$  and  $\mathfrak{K}_{2a}$  is taken to be the third lowest s solution to  $H_{2a}$  rather than the second lowest. As would be expected, this process works well, and the wave functions converge as well as for the ground state. We could also use the same process with (4) where we would take the first and third solutions of  $H_A$  to be occupied. However, in this case the second and third s-type solutions are nearly degenerate, and the iteration process diverges (this problem also occurs for the  $3^2P$  state of Li). Thus we were forced to use (1) for

TABLE II. The energies for the 2s orbitals and the virtual orbitals of Li from UHF and GF calculations (the basis set was based on the effective quantum numbers,  $n^*$ , from the term values of the lower s and p states. We used 1s, 2s=2.7; 2s, 3s=0.63; 3s, 4s=0.39; 4s, 5s=0.28; 5s=0.22; 2p, 3p=0.51; 3p, 4p=0.34; 4p, 5p=0.25 where the Slater orbital is represented by nl and the number is the orbital exponent) on the ground state,  ${}^2S[(1s)^2(2s)]$ . Energies are in hartrees.

	UH	$\mathbf{F}$	•	GF		
State	H <sub>2a</sub> Different <sup>C</sup> Hamiltonian	$H_{A}$ Same <sup>d</sup> Hamiltonian	H <sub>2a</sub> Different <sup>c</sup> Hamiltonian	H <sub>A</sub> Same <sup>d</sup> Hamiltonian	SCF <sup>a</sup>	Experimental <sup>b</sup>
2s	-0.1961	-0.1961	-0.1963	-0.1963	-0.1963	-0.1982
3s	-0.0738	+0.0017	-0.0738	-0.0152	-0.0738	-0.0742
4s	-0.0385	+0.0077	-0.0385	-0.0060		-0.0386
5s	-0.0236	+0.0214	-0.0236	+0.0033		-0.0236
6s	-0.0123	+0.0548	-0.0123	+0.0289	• • •	-0.0159
2 <i>p</i>	-0.1283	+0.0036	-0.1283	-0.0273	-0.1287	-0.1302
3p	-0.0567	+0.0106	-0.0567	-0,0084	-0.0568	-0.0572
4p	-0.0317	+0.0230	-0.0317	+0.0005	• • •	-0.0320
5 <i>p</i>	-0.0193	+0.0516	-0.0193	+0.0221	• • •	-0.0204

<sup>&</sup>lt;sup>a</sup>To this many places the UHF and GF methods yield the same orbital energies.

calculations on these excited states of Li.
All these results are exactly the same in the GF case. The wave function is taken as

$$G_f(\phi_{1a}\phi_{2a}\phi_{1b}\alpha\alpha\beta),$$

and the equations for the optimum orbitals are as in (1),  $^{7,9}$  where the forms of  $H_{1a}$ ,  $H_{2a}$ , and  $H_{1b}$  are more complicated than (2). For example, after making a Roothaan expansion of the orbitals in terms of basis functions  $\{X_{1a}\}$  the matrix equivalent of  $H_{2a}$  has the form  $^{7,9,23}$ .

$$\begin{split} H_{\mu\nu}^{\quad 2a} &= (1 + \frac{1}{2}\langle 1a \mid 1b\rangle^2)\langle \mu \mid h \mid \nu\rangle \\ &+ \frac{1}{2}\langle \langle \mu \mid h \mid 1b\rangle - \langle \mu \mid h \mid 1a\rangle\langle 1a \mid 1b\rangle)\langle 1b \mid \nu\rangle \\ &+ \frac{1}{2}\langle \mu \mid 1b\rangle\langle\langle 1b \mid h \mid \nu\rangle - \langle 1a \mid h \mid \nu\rangle\langle 1b \mid 1a\rangle) \\ &+ \frac{1}{2}\langle \mu \mid 1b\rangle\langle 1b \mid \nu\rangle\langle\langle 1a \mid h \mid 1a\rangle - E) \\ &+ (\mu\nu \mid 1a1a) + (\mu\nu \mid 1b1b) + (\mu\nu \mid 1a1b)\langle 1a \mid 1b\rangle \\ &- (\mu 1a \mid \nu 1a) + \frac{1}{2}(\mu 1b \mid \nu 1b) \\ &- \frac{1}{2}\langle 1a \mid 1b\rangle[(\mu 1a \mid \nu 1b) + (\mu 1b \mid \nu 1a)] \\ &+ \frac{1}{2}[(\mu 1b \mid 1a1a) - (\mu 1a \mid 1a1b)]\langle 1b \mid \nu\rangle \end{split}$$

 $+\frac{1}{2}\langle \mu | 1b \rangle [(1a1a | \nu 1b) - (1a1b | \nu 1a)].$ 

Again we can add terms to  $H_{1a}$  and  $H_{2a}$  to get an  $H_A$  such that  $\phi_{1a}$  and  $\phi_{2a}$  are both solutions of  $H_A$  as in (4). For example, we would add

$$\Delta H \frac{2a}{\mu \nu} = \frac{1}{2} \langle 2a | 1b \rangle^{2} \langle \mu | h | \nu \rangle$$

$$- \frac{1}{2} \langle \mu | h | 2a \rangle \langle 2a | 1b \rangle \langle 1b | \nu \rangle$$

$$- \frac{1}{2} \langle \mu | 1b \rangle \langle 1b | 2a \rangle \langle 2a | h | \nu \rangle$$

$$+ \frac{1}{2} \langle \mu | 1b \rangle \langle 1b | \nu \rangle \langle 2a | h | 2a \rangle$$

$$+ (\mu \nu | 2a2a) + (\mu \nu | 2a1b) \langle 2a | 1b \rangle$$

$$- (\mu 2a | \nu 2a)$$

$$- \frac{1}{2} \langle 2a | 1b \rangle [(\mu 2a | \nu 1b) + (\mu 1b | \nu 2a)]$$

$$+ \frac{1}{2} [(\mu 1b | 2a2a) - (\mu 2a | 2a1b)] \langle 1b | \nu \rangle$$

$$+ \frac{1}{2} \langle \mu | 1b \rangle [(2a2a | \nu 1b) - (2a1b | \nu 2a)]$$

to  $H_{\mu\nu}^{\phantom{\mu\nu}2a}$  to get  $H_{\mu\nu}^{\phantom{\mu\nu}A}$ . (Note that  $\sum_{\nu}\Delta H_{\mu\nu}^{\phantom{\mu\nu}2a}$   $\times \langle \nu \, | \, 2a \rangle = 0$ .) Just as in the UHF case, we find that the virtual solutions of  $H_{2a}$  are good approximations to the excited orbitals, whereas the virtual solutions to  $H_A$  are not (see Table II).

Of course the above results apply equally well to other symmetries and other systems of various numbers of electrons.

bExperimental excitation energies from C. Moore, footnote c to Table I. This column is not strictly comparable with the other columns because of differences in the core states and correlation energy for the various states. However, these differences are small for the excited states.

 $<sup>^{</sup>m c}$ No self-repulsion terms are included in the one-electron Hamiltonian. Thus each electron has a different Hamiltonian.  $^{
m d}$ Self-repulsion terms are included in the one-electron Hamiltonian. Thus  $\phi_{1a}$  and  $\phi_{2a}$  are both eigenstates of the same Hamiltonian.

### DEPENDENCE OF PROPERTIES UPON EXCITED STATES

The total electronic density at the nucleus,  $\rho(0)$ , is given in Table III for the various states of Li. It is interesting to note that for ns states the  $\rho(0)$ decreases as we go to large n, eventually decreasing to the  $\rho(0)$  of Li<sup>+</sup>, but that for np and nd states  $\rho(0)$  is less than the  $\rho(0)$  of Li + and increases to  $\rho(0)$  of Li + as n and l increase. 24 An explanation of this behavior is that the valence electron, nl, partly shields the nucleus from the core electrons and thus leads to a slightly lower contribution to ho(0) from the core electrons,  $ho(0)_{\hbox{core}}$ , as compared to  $Li^+$ . If the added electron is in an nsstate, it has a nonzero amplitude at the nucleus which leads to an increase in  $\rho(0)$  that more than cancels the previous decrease. But for  $l \neq 0$ there is no compensating term, and  $\rho(0)$  is lower for the atom than for Li  $^+$ . (Note that  $ho(0)_{\hbox{core}}$  is nearly independent of l for ns and np states.)

In addition, (see Table I) we have the order 2s < 2p < 3s < 3p < 3d for total energy, whereas for  $\langle \sum r^2 \rangle$  we have the order 2s < 2p < 3s < 3d < 3p. Thus using  $\langle \sum r^2 \rangle$  as the criterion,  $\phi_{3d}$  is less diffuse than  $\phi_{3p}$ , but more diffuse than  $\phi_{3s}$ . However, for  $\langle 1/r^3 \rangle_{nl}$  the order is 2p < 3p < 3d, which might be expected since the inner loop of the 3p should count heavily here.

The value of  $\langle 1/r^3 \rangle$  for various states nl of the same l are often<sup>25,26</sup> expected to scale as  $(1/n^*)^3$ , where  $n^*$  is the effective quantum to use in the hydrogen atom formula in order to obtain the term energy. It has been less clear how the spin density term would scale, although it has also been assumed<sup>26</sup> to scale as  $(1/n^*)^3$ . In Table IV we compare the scaling of the hyperfine constants to that of  $n^*$ . We see that  $(1/n^*)^3$  leads to results too small by 1% and 8% for Q(0) of the  $3^2S$  and  $3^2P$  states and too small by 2%, 4%, and 13% for  $\langle 1/r^3\rangle$ ,  $a_{1/2}$  and  $a_{3/2}$  of the  $3^2P$  state.

### CONCLUSIONS

For the excited states of Li as for the ground state, the GF wave function leads to good values for the hyperfine parameters. Thus the major error in the Hartree-Fock value of the hyperfine splitting constants of Li is due to the lack of core polarization.

TABLE III. The density at the nucleus,  $\rho(0)$ , for Li<sup>+</sup> and several states of Li.  $\rho(0)_{\mbox{core}}$  is the density at the nucleus due to the 1s electrons. The Hartree-Fock wave functions of Weiss (see Ref. 4) for Li and Roothaan, Sachs, and Weiss [Rev. Mod. Phys. 32, 186 (1960)] for Li<sup>+</sup> were used.

`	ρ(0)	ho(0)
Li 2 <sup>2</sup> S	13.816	13,652
$3^2S$	13.708	13.666
$2^2P$	13.653	13.653
$3^2P$	13.666	13.666
$3^2D$	13.672	13.672
Li <sup>+</sup> 1 <sup>1</sup> S	13.674	13.674

The calculated  $a_{3/2}$  from GF and CI calculations disagree with the values deduced from the level-crossing experiments. This casts doubt on the interpretation of these experiments in terms of effective  $\langle 1/r^3 \rangle$  and Q(0) that are independent of magnetic field.

The use of different Hamiltonians for different orbitals leads to improved convergence of excited-state wave functions and also leads to virtual orbitals closely related to the orbitals of other excited states.

## APPENDIX A. THE CALCULATED WAVE FUNCTIONS

The calculated orbitals for the UHF and GF wave functions are available upon request. The basis functions (which are Slater orbitals), the orbital energies  $(\epsilon_i)$ , and the amplitude at the nucleus  $[\phi_i(0)]$  are given in Table V.

### APPENDIX B: HYPERFINE SPLITTING CONSTANTS

We define the hyperfine splitting constant,  $A_J$ , as the proportionality constant between  $\langle \vec{\mathbf{I}} \cdot \vec{\mathbf{J}} \rangle$  and the perturbation energy  $E = A_J \langle \vec{\mathbf{I}} \cdot \vec{\mathbf{J}} \rangle$ . Then we define  $a_J$  by  $A_J = g_S \mu_B g_N \mu_N a_J$ , where  $\mu_B$  and  $\mu_N$  are the Bohr and nuclear magnetons, and  $g_S$  and  $g_N$  are the gyromagnetic ratios for the electron and nucleus. If the different states are all taken to be eigenstates of  $\hat{L}^2$  and  $\hat{S}^2$ , then  $a_J$  is  $a_J$ ?

TABLE IV. The scaling of the hyperfine constants for the ns and np states of Li (n=2,3).  $n^*$  is the effective quantum number obtained from the term energies (from Moore, footnote c to Table I).

	n*	$(1/n^*)^3$	Q(0)	$\langle 1/r^3  angle$	$a_{1/2}$	$a_{3/2}$
2s	1.588 <sup>a</sup>	•••	0.2406			
3 <i>s</i>	$2.596^{a}$	• • •	0.056 22		****	
ratio		0.229	0.234	• • •	• • •	• • • •
2p	1.966 <sup>a</sup>		-0.023 04	0.05861	0.2206	-0.033 13
3 <i>p</i>	2.956 <sup>a</sup>	• • •	-0.007318	0.01760	0.06735	-0.01106
ratio		0.294	0.318	0.300	0.305	0.334

TABLE V. The orbital energy and amplitude at the nucleus from the UHF and GF wave functions of Li. All quantities are in Hartree atomic units. Basis sets: the entries represent the  $nl\zeta$  of each Slater orbital (see Ref. 10).

			Orbital Energy		Amplitude at the Nucleus				
State		$\epsilon_{1a}$	$\epsilon_{1b}$	$\epsilon_{2a}$	$\phi_{1a}^{(0)}$	$\phi_{1b}^{(0)}$	$\phi_{2a}^{(0)}$		
$2^2 S^{a}$	UHF	-2.48670	-2.46873	-0.19637	2.61848	2.60684	-0.40486		
	$_{ m GF}$	-2.498 75	-2.46388	-0.19649	2.63388	2.59173	-0.401 95		
$3^2 s^a$	UHF	-2.66484	-2.66074	-0.07380	2.61554	2.61287	-0.19641		
	$_{ m GF}$	-2.66766	-2.65952	-0.07380	2.61922	2,60920	-0.19612		
$2^2 P^{\mathbf{b}}$	UHF	-2.53133	-2.53005	-0.128 68	2.611 13	2.61448	0.00000		
	GF	-2,530 55	-2.53220	-0.12872	2.60614	2.61941	0.00000		
$3^2P^{\mathbf{C}}$	UHF	-2.677 80	-2.67735	-0.05678	2.61349	2.61454	0.00000		
	$_{ m GF}$	-2.677 55	-2.677 99	-0.05678	2.61191	1.61611	0.00000		
$3^2 D^{\mathrm{d}}$	UHF	-2.681 23	-2.68123	-0.05556	2.614 59	2.61460	0.00000		
	GF	-2.681 22	-2.68124	-0.05556	2.61458	2.61461	0.00000		

a<sub>1s</sub> 3.0, 3s 8.7, 3s 3.398, 3s 2.544, 3s 1.24, 3s 0.757, 3s 0.345, 4s 0.345.

$$\begin{split} a_{J} = & \left(\frac{2}{g_{s}}\right) \! a_{I} \! \frac{\langle \vec{\mathbf{L}} \cdot \vec{\mathbf{J}} \rangle}{J(J+1)} + \frac{8\pi}{3} \, a_{C} \! \frac{\langle \vec{\mathbf{S}} \cdot \vec{\mathbf{J}} \rangle}{J(J+1)} \\ & + a_{d} \! \frac{3 \langle \vec{\mathbf{S}} \cdot \vec{\mathbf{L}} \rangle \langle \vec{\mathbf{L}} \cdot \vec{\mathbf{J}} \rangle - L(L+1) \langle \vec{\mathbf{S}} \cdot \vec{\mathbf{J}} \rangle}{J(J+1)} \quad . \end{split}$$

Relativistic effects and breakdown of LS coupling have been ignored and  $g_{\tilde{l}}=1$  has been assumed. The constants in the above equation are defined as

$$\begin{split} a_l &= \frac{1}{L} \left\langle LSLS \left| \sum_e \hat{\frac{l}{r^3}} \right| LSLS \right\rangle, \\ a_c &= \frac{1}{S} \left\langle LSLS \left| \sum_e \delta(r) \hat{s}_z \right| LSLS \right\rangle, \\ a_d &= \frac{1}{SL(2L-1)} \left\langle LSLS \left| \sum_e \left( \frac{3z^2 - r^2}{r^5} \right) \hat{s}_z \right| LSLS \right\rangle, \\ \left\langle \vec{\mathbf{L}} \cdot \vec{\mathbf{J}} \right\rangle &= \frac{1}{2} \left[ J \left( J + 1 \right) + L \left( L + 1 \right) - S \left( S + 1 \right) \right], \\ \left\langle \vec{\mathbf{S}} \cdot \vec{\mathbf{J}} \right\rangle &= \frac{1}{2} \left[ J \left( J + 1 \right) + S \left( S + 1 \right) - L \left( L + 1 \right) \right], \\ \left\langle \vec{\mathbf{S}} \cdot \vec{\mathbf{L}} \right\rangle &= \frac{1}{2} \left[ J \left( J + 1 \right) - S \left( S + 1 \right) - L \left( L + 1 \right) \right], \end{split}$$

where  $|LSM_LM_S\rangle$  denotes a many-electron wave function having quantum numbers L, S,  $M_L$ , and  $M_S$ . The subscripts l,c, and d refer to orbital, Fermi contact, and spin-dipolar contributions, respectively.

For the HF, PUHF, and GF wave functions of Li,  $a_I$  and  $a_d$  simplify since

$$\left\langle LSLS \middle| \sum_{e} \frac{\hat{l}_z}{r^3} \middle| LSLS \right\rangle = \left\langle \frac{1}{r^3} \right\rangle$$

and

$$\frac{1}{S} \left\langle LSLS \left| \sum_{e} \frac{(3z^2 - r^2)}{r^5} \hat{s}_z \right| LSLS \right\rangle = \frac{-2l}{(2l+3)} \left\langle \frac{1}{\gamma^3} \right\rangle,$$

where  $\langle 1/r^3 \rangle$  is a one-electron radial integral over the  $l \neq 0$  orbital (with  $m_l = l$ ).

Using the values  $g_{\rm S}=2.002\,3192,~\mu_B=9.2732$   $\times 10^{-24}$  Am²,  $\mu_N=5.050\,50\times 10^{-27}$  Am², 1 eV = 1.602  $10\times 10^{-19}$ J, 1eV = 2.418 04  $\times 10^{14}$ Hz, and  $a_0=0.529\,167$  Å, we find that to convert  $a_J$  from atomic units to Mc/sec we have to multiply by  $c=95.51975g_N$ . Thus for Li<sup>7</sup> using  $g_N=3.256\,31/1.5$ , we obtain  $c({\rm Li}^7)=207.36$ .

Using the above equations and letting  $a_C \equiv Q(0)$  we obtain

for  $ns^2S$  states:  $a_{1/2} = 8.3776Q(0)$ ,

for  $np^2P$  states:  $a_{1/2} = 2.6651\langle 1/r^3 \rangle - 2.7925Q(0)$ .

 $a_{3/2} = 0.53256\langle 1/r^3 \rangle + 2.7925Q(0),$ 

for  $nd^2D$  states:  $a_{3/2} = 0.99930 \langle 1/r^3 \rangle - 1.6755Q(0)$ ,

 $a_{5/2} = 0.28525\langle 1/r^3 \rangle + 1.6755Q(0).$ 

In the derivation it is necessary to assume that the wave functions are eigenfunctions of  $L^2$  and  $S^2$  and can be combined into eigenfunctions of  $J^2$ . This is correct for the HF, PUHF, and GF wave functions for Li, but not for the UHF wave functions. Thus in the UHF case it is not correct to use the above equations (or any others for that matter) for  $a_J$ , since the UHF wave functions are not eigenfunctions of  $S^2$  or  $J^2$ . However, we will follow the usual practice of using the above equations for UHF also.

 $b_{1s}$  3.0, 3s 8.7, 3s 3.398, 3s 3.544, 2p 1.5, 4p 2.12, 4p 1.275, 4p 0.785, 4p 0.566.

<sup>&</sup>lt;sup>c</sup>Same as b except that 4p 0.566 is replaced by 4p 0.63, 4p 0.37, 4p 0.21.

 $d_s$  orbtials as in b, 3d 1.0, 5d 0.98, 5d 0.60, 5d 0.404, 5d 1.762.

## APPENDIX C: UPPER BOUNDS ON THE ENERGY FOR EXCITED STATES

The UHF and GF wave functions for the excited states of Li lead to upper bounds on the exact energy, even though the wave functions are not orthogonal to the ground state. Below we sketch the proof of this for GF wave functions. 30 A sim ilar theorem for HF wave functions has been proved using a different approach by Perkins. 31

Let 
$$\psi_1 = G_f(\phi_{1s}^a \phi_{2s}^a \phi_{1s}^b \alpha \alpha \beta),$$
 (C-1) 
$$\psi_2 = G_f(\phi_{1s}^a \phi_{3s}^a \phi_{1s}^b \alpha \alpha \beta),$$

and consider linear combinations

$$\psi = C_1 \psi_1 + C_2 \psi_2 \tag{C-2}$$

of these functions. If the coefficients in (C-2) are chosen to make the total energy stationary, we find that the coefficients for the ith solution satisfy

$$\sum_{k} \langle \psi_{j} | H - E_{i} | \psi_{k} \rangle C_{j} = 0. \tag{C-3}$$

In addition, from the Hylleraas-Undheim-Mac Donald theorem<sup>32</sup> the calculated energy of the ith solution is an upper bound on the exact energy for the ith solution allowed by the symmetries of (C-1). But if the orbitals in (C-1) are solutions of the GF equations corresponding to (1), where  $\phi_{1s}a$ ,  $\phi_{1s}b$ , and  $\phi_{3s}^{a}$  are used in the field terms, then from the Brillouin theorem for GF wave functions, 33 we have that

$$\begin{split} &\langle \phi_{1s}^{\phantom{1}s} \phi_{c} \phi_{1s}^{\phantom{1}b} \mid_{H-E_{2} \mid O_{ff} \phi_{1s}^{\phantom{1}a} \phi_{3s}^{\phantom{3}a} \phi_{1s}^{\phantom{1}b} \rangle \\ &= \langle \phi_{c} \mid_{H_{2a}}^{\phantom{2}GF} \mid \phi_{3s}^{\phantom{3}a} \rangle = \epsilon_{3s}^{\phantom{3}a} \langle \phi_{c} \mid \phi_{3s}^{\phantom{3}a} \rangle = 0 \end{split} \tag{C-4}$$

for any  $\phi_C$  orthogonal to  $\phi_{3S}{}^a$ . Thus the solutions of (C-3) are  $C_1$ =1,  $C_2$ =0 and  $C_1$ =0,  $C_2$ =1; that is, the  $\psi_1$  and  $\psi_2$  in (C-1) are solutions of (C-3). But<sup>32</sup>

$$E_1 = \langle \psi_1 | H | \psi_1 \rangle / \langle \psi_1 | \psi_1 \rangle = E^* + \epsilon_{2s}^a,$$

$$E_2 = \langle \psi_2 | H | \psi_2 \rangle / \langle \psi_2 | \psi_2 \rangle = E^* + \epsilon_{3s}^a$$

where  $E^*$  is the part of the energy term which depends only on  $\phi_{1a}$  and  $\phi_{1b}$ ,  $\epsilon_{3s}{}^a$  is the eigenvalue for the occupied orbital  $\phi_{3s}{}^a$ , and  $\epsilon_{2s}{}^a$  is the eigenvalue of the virtual orbital  $\phi_{2s}{}^a$ . Thus if  $\epsilon_{2s}{}^a < \epsilon_{3s}{}^a$ , then  $E_1 < E_2$  and  $\psi_2$ , the GF solution for the  $3^2S$  state of Li, yields an upper bound on the exact energy of the 32S state. This occurs despite the fact that  $\psi_2$  is not orthogonal to the exact GF solution for the  $2^2S$  state and a fortiori is not orthogonal to the exact solution of the 22S state.

Now consider the case where the N-electron wave function  $\psi_n$  is a self-consistent GF solution of the with state of a given symmetry and  $\psi_j(j < n)$  are constructed by replacing  $\phi_{nS}{}^a$  by  $\phi_{jS}{}^a$ , a lower energy virtual orbital of  $H_{2a}$ . Then (C-4) still holds for  $\phi_C = \phi_{jS}{}^a$  if  $j \neq n$ , and  $\psi_n$  is already a solution of  $H_{2a}$ . solution of the equations corresponding to (C-3). Thus if there are n-1 virtual orbitals of lower energy, then  $E_n=\langle \psi_n \mid H \mid \psi_n \rangle/\langle \psi_n \mid H \mid \psi_n \rangle$  is an upper bound on the energy of the nth state of this symmetry. This proof does not depend on the symmetry of the valence orbitals; it applies only to excited states obtained by exciting a single electron.

<sup>11</sup>For the 2 <sup>2</sup>S state of Li the CI calculations quoted (Ref. 12) used interelectronic coordinates. The best CI calculation on this state not using interelectronic coordinates is that of Weiss (Ref. 3) which yielded a poor value (Ref. 13) for Q(0) of 0.2065 [compared to the experimental value of 0.2313 (Ref. 14)]. The reason for this poor Q(0) is that Weiss did not include any configurations combining s-type core orbitals into a triplet. It is that such configurations which are needed for the core orbitals to contribute to Q(0). The CI calculation on the 2 2P state (Weiss, Ref. 4) did include such configurations, and thus can be expected to yield a more reliable value for Q(0). However, it is possible that an insufficient number of such configurations was used; in this case the CI value of Q(0) may not have converged to the correct value.

<sup>12</sup>G. J. Ritter, Canadian J. Phys. 43, 770 (1965). <sup>13</sup>K. C. Brog, T. G. Eck, and H. Wieder, Phys. Rev. 153, 91 (1967).

<sup>14</sup>In order to interpret the level crossing experiments in terms of  $\langle 1/r^3 \rangle$  and Q(0), Brog et al. (Ref. 16) allowed a general mixing of zero-field configurations by the external field. But they allowed only the six zero-field configurations appropriate for Li 2 2P, and assumed that these were related to each other and to

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<sup>‡</sup>Contribution No. 3703

<sup>&</sup>lt;sup>1</sup>L. W. Anderson, F. M. Pipkin, and J. C. Baird, Phys. Rev. 116, 87 (1959).

<sup>&</sup>lt;sup>2</sup>J. S. M. Harvey, Proc. Roy. Soc. (London) <u>A285</u>, 581

<sup>&</sup>lt;sup>3</sup>A. W. Weiss, Phys. Rev. 122, 1826 (1961).

<sup>&</sup>lt;sup>4</sup>A. W. Weiss, Astrophys. J. <u>138</u>, 1262 (1963).

<sup>&</sup>lt;sup>5</sup>J. A. Pople and R. K. Nesbet, J. Chem. Phys. <u>22</u>, 571 (1954); J. C. Slater, Phys. Rev. 82, 538 (1951); G. W. Pratt, Jr., Phys. Rev. 102, 1303 (1956); R. K. Nesbet, Proc. Roy. Soc. (London) A230, 312 (1955); A. T. Amos and G. G. Hall, ibid. A263, 483

<sup>&</sup>lt;sup>6</sup>P. O. Löwdin, Advan. Chem. Phys. <u>2</u>, 207 (1959); S. M. Blinder, Advan. Quant. Chem. 2, 47 (1965).

<sup>&</sup>lt;sup>7</sup>W. A. Goddard, III, Phys. Rev. <u>157</u>, 81 (1967). <sup>8</sup>W. A. Goddard, III, Phys. Rev. <u>157</u>, 93 (1967).

<sup>&</sup>lt;sup>9</sup>W. A. Goddard, III, J. Chem. Phys. <u>48</u>, 450 (1968).

<sup>10</sup>A Slater orbital is a function of the form  $x_{nlm} = N_n r^{n-1} e^{-\xi r} Y_{lm}$ , where  $\xi$  is a variable parameter called the orbital exponent and  $N_n$  is a normalization factor.

 $\langle 1/r^3\rangle$  and Q(0) as they would be for Hartree-Fock wave functions.

<sup>15</sup>R. W. B. Ardill and A. L. Stewart, J. Chem. Phys. <u>47</u>, 4853 (1967); Proc. Phys. Soc. (London) <u>92</u>, 296 (1967).

<sup>16</sup>D. A. Goodings, Phys. Rev. 123, 1706 (1961).

<sup>17</sup>M. H. Cohen, D. A. Goodings, and V. Heine, Proc. Phys. Soc. (London) <u>73</u>, 811 (1959).

<sup>18</sup>G. D. Gaspari, W. M. Shyu, and T. P. Das, Phys. Rev. 134, A852 (1964).

<sup>19</sup>W. Marshall, J. Phys. Soc. Japan <u>17</u>, Suppl. B-120 (1962).

<sup>20</sup>Burke (Ref. 12b) has also concluded that the concept of core polarization by non-s electrons is valid, but Bergreen and Wood [Phys. Rev. <u>130</u>, 198 (1963)] question the idea of core polarization by s electrons.

 $^{21}$ J. B. Martin and A. W. Weiss, J. Chem. Phys.  $\underline{39}$ , 1618 (1963).

<sup>22</sup>T. Koopmans, Physica <u>1</u>, 104 (1933).

 $^{23}$ Here the orbitals are taken to be real,  $\langle 1a | 1b \rangle$  denotes an overlap integral,  $\langle 1a | h | 1b \rangle$  denotes a one-electron integral, and

$$\begin{split} (\alpha\beta\mid\gamma\delta) &= \int\! d\vec{\mathbf{x}}_1 \phi_\alpha^{\ *}(1)\,\phi_\beta^{\ }(1) \\ &\times \int\! d\vec{\mathbf{x}}_2^{\ }(1/r_{12})\phi_\gamma^{\ *}(2)\,\phi_\delta^{\ }(r) \end{split}$$

is the usual notation for a two-electron integral. We have multiplied (4) by  $(1+\frac{1}{2}\langle 1a|\ 1b\rangle^2)$  for convenience

<sup>24</sup>Comparing the  $\rho(0)$  for the HF wave function of Li<sup>+</sup> to the value for the very accurate Pekeris wave function [Phys. Rev. 126, 143 (1962)],  $\rho(0) = 13.704$ , we would expect the  $\rho(0)$  for the various states of Li in Table II to also be lower than the exact values by about 0.03 or 0.2%.

<sup>25</sup>R. Isler, S. Marcus, and R. Novick, Bull. Am. Phys. Soc. 11, 62 (1966).

<sup>26</sup>B. Budick, H. Bucka, R. J. Goshen, A. Landman, and R. Novick, Phys. Rev. <u>147</u>, 1 (1966).

<sup>27</sup>For example, see R. E. Trees, Phys. Rev. <u>92</u>, 308 (1953).

<sup>28</sup>E. R. Cohen and J. W. M. Dumond, Rev. Mod. Phys. <u>37</u>, 537 (1965).

<sup>29</sup>N. F. Ramsey, <u>Molecular Beams</u> (Oxford, Clarendon Press, London, 1956) p. 172.

<sup>30</sup>The basic idea of using a configuration interaction approach to prove this theorem was suggested by Professor R. M. Pitzer, private communication, 1966.

<sup>31</sup>J. F. Perkins, J. Chem. Phys. <u>42</u>, 3927 (1965).
 <sup>32</sup>E. Hylleraas and B. Undheim, Z. Physik <u>65</u>, 759 (1930); J. K. L. MacDonald, Phys. Rev. <u>43</u>, 830 (1933).

 $^{33}$ W. A. Goddard, III, J. Chem. Phys.  $\underline{48}$ , 5337 (1968).

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### Relativistic Calculations of Electron Binding Energies by a Modified Hartree-Fock-Slater Method

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Relativistic atomic wave functions and electron binding energies have been calculated by a modified Hartree-Fock-Slater method. The local exchange approximation, originally introduced by Slater, is modified in such a way that the total energy of the system is minimized. Approximate expressions for the optimum exchange potential are suggested for all atoms and ions. The total atomic energies and electron binding energies obtained by this method agree extremely well with the corresponding Hartree-Fock results in the cases where such data are available. Various corrections to the theoretical binding energies are discussed, in particular the effect of rearrangement during the ionization process. It is found that this effect is of importance for inner shells in all elements and is responsible for the main discrepancy between experimental and previous theoretical results for light and medium-heavy elements. For heavy elements other effects are of importance, and various possible sources of the residual discrepancy between theory and experiments are discussed.

### I. INTRODUCTION

During the last decade our knowledge of electron binding energies in atoms has increased considerably, due in particular to the development of the electron-spectroscopic method (ESCA) by Siegbahn  $et\ al.^1$  This method has been applied to almost all elements between lithium and berkelium and to inner as well as outer shells. The accuracy of the

new data is so high that improved theoretical calculations are justified. The electron binding energies are often used to check the accuracy of electronic wave functions. When the accuracy of the calculations increases, several effects have to be taken into account. Besides the relativistic effects, which of course are dominant for heavy elements, it is necessary to consider relaxation mechanisms, correlation effects, core polariza-