

Note on space-times that admit constant electromagnetic fields*

Douglas M. Eardley

California Institute of Technology, Pasadena, California 91109
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All space-times that admit a covariantly constant, test, electromagnetic field are constructed. All solutions to the Einstein-Maxwell equations with constant electromagnetic field are given.

1. INTRODUCTION

Honig, Schücking, and Vishveshwara recently gave an elegant discussion¹ of the motion of a charged test particle in any space-time (M, g) with a covariantly constant (hence source-free) electromagnetic field F . But they did not address the question: Which (M, g) admit such an F , even a test field F ? This note answers this question. The argument will be familiar to mathematicians, but physicists may find it novel. For simplicity, we will confine the work to some local neighborhood in (M, g) .

2. HOLONOMY OF SPACE-TIME

Choose any point $p \in M$ in space-time. Fix, once and for all, a simply connected neighborhood U of p . Parallely carry an orthonormal tetrad O_p around any closed path in U beginning and ending at p , yielding some new tetrad O'_p : $O'_p = \Lambda O_p$, where Λ is some homogeneous Lorentz transformation depending only on the choice of path. The set of all such Λ at p for all possible paths forms Cartan's (local) holonomy group^{2,3,4} $H(p)$. $H(p)$ is a subgroup of the homogeneous Lorentz group L ; $H(p)$ is in fact independent of $p \in U$. Roughly, the higher the symmetry of (M, g) , the smaller $H(p)$ is.

Let T be any geometric-object field on (M, g) for which covariant differentiation ∇ is defined. Can there exist in U a covariantly constant T , $\nabla T = 0$? If so, T can be uniquely constructed by giving its value $T(p)$ at one point $p \in U$, and then carrying T parallely all over U along paths in U . This construction must be path-independent; equivalently, $T(p)$ must be left invariant at p when it is carried parallely around any closed path beginning and ending at p . That is, $T(p)$ must be invariant under the action of $H(p)$:

Lemma (see e.g., Schouten²): (M, g) (locally) admits a covariantly constant, test, field T iff there exists $T(p)$ at any point $p \in M$, invariant under the holonomy group $H(p)$.

3. CONSTANT, TEST, ELECTROMAGNETIC FIELD

The differential problem, "solve $\nabla F = 0$," is thus reduced to an algebraic problem (Cartan's favorite trick!): Choose a point p ; find an electromagnetic field tensor $F(p) \neq 0$ such that its invariance group $G[F(p)]$,

$$G[F(p)] \equiv \left\{ \Lambda \in L \mid \Lambda F(p) = F(p) \right\},$$

contains the holonomy group: $H(p) \subseteq G[F(p)] \subseteq L$.

Choose a favored observer at p with tetrad $O_p = (e_{\hat{t}}, e_{\hat{x}}, e_{\hat{y}}, e_{\hat{z}})$ so that the tetrad components $F_{\hat{\mu}\hat{\nu}}$ of $F(p)$ reduce to one of two canonical forms⁵: " $F(p)$ null" or " $F(p)$ nonnull."

(A) $F(p)$ null: $F_{\hat{t}\hat{x}} = -F_{\hat{x}\hat{t}} = F_{\hat{t}\hat{z}} = -F_{\hat{z}\hat{t}} = A = \text{const}$, other components vanish. $G[F(p)]$ is generated by the two null rotations⁶ which leave invariant the null direc-

tion defined by $k = e_{\hat{t}} + e_{\hat{z}}$. Define the spinor⁷ $o^A(p)$ by $o^A(p)\bar{o}^{\dot{A}}(p) = k^{A\dot{A}}$; then $G[o^A(p)] = G[F(p)]$. Therefore (by lemma) (M, g) must admit a covariantly constant spinor field o^A (equivalently, a covariantly constant, complex, null bivector field $F + i^*F$). By a result of Ehlers and Kundt,⁸ it is necessary and sufficient that in some local coordinates

$$g = 2K(u, x, y)du^2 + 2dudv + dx^2 + dy^2, \quad (1a)$$

and

$$F = 2^{1/2}Adu \wedge dx. \quad (1b)$$

(B) $F(p)$ nonnull: $F_{\hat{t}\hat{z}} = -F_{\hat{z}\hat{t}} = A \cos\theta = \text{const}$, $F_{\hat{t}\hat{x}} = -F_{\hat{x}\hat{t}} = A \sin\theta = \text{const}$, other components vanish. $G[F(p)]$ is the direct product of the one-parameter group of boosts in the (tz) plane with the commuting one-parameter group of rotations in the (xy) plane. It is a fundamental result³ that if the tangent space T_p is reducible under $H(p)$, then (M, g) is correspondingly reducible into the direct product of (pseudo-) Riemannian manifolds of lower dimension. Here, T_p reduces to the (tz) and (xy) planes under $H(p)$, so $(M, g) = (M_+, g_+) \otimes (M_-, g_-)$, where (M_+, g_+) is a Lorentzian 2-manifold and (M_-, g_-) is a Riemannian 2-manifold. The vector fields $e_{\hat{t}}$ and $e_{\hat{z}}$ lie entirely in (M_+, g_+) ; $e_{\hat{x}}$ and $e_{\hat{y}}$ lie entirely in (M_-, g_-) . In some coordinates,

$$g = g_+ + g_-, \quad (2a)$$

where

$$g_+ = g_{+ab}(t, z)dx^a dx^b, \quad x^a \equiv (t, z), \quad (2b)$$

$$g_- = g_{-ij}(x, y)dx^i dx^j, \quad x^i \equiv (x, y), \quad (2c)$$

and

$$F = A \cos\theta(-g_+)^{1/2}dt \wedge dz + A \sin\theta(g_-)^{1/2}dx \wedge dy. \quad (2d)$$

Equations (1) and (2) give all space-times (M, g) , and all test fields F , that solve $\nabla F = 0$.

4. ELECTROVAC SOLUTIONS

Now impose the Einstein-Maxwell equations for a nontest F , to find all solutions with covariantly constant F . The resulting electrovac space-times are well known.

(A) F null: The Einstein-Maxwell equations for Eqs. (1) read

$$(\partial_x^2 + \partial_y^2)K(u, x, y) = -4A^2. \quad (3)$$

The general solution K is a linear superposition, $K = K_{\text{em}} + K_{\text{grav}}$, where K_{em} is the particular solution,

$$K_{\text{em}} = -A^2(x^2 + y^2),$$

and K_{grav} is any homogeneous solution,

$$(\partial_x^2 + \partial_y^2)K_{\text{grav}}(u, x, y) = 0.$$

K_{em} represents the transverse, "monopole," always focusing gravitational disturbance due to \mathbf{F} ; this disturbance is homogeneous for local observers (despite the dependence of K_{em} on x and y). K_{grav} represents an arbitrary "plane-fronted gravitational wave with parallel rays"⁸ ("pp wave"). So the electrovac space-time given by Eqs. (1) and (3) describes an arbitrary, gravitational pp wave traversing a region of constant, null \mathbf{F} , such that the wave direction is everywhere parallel to the Poynting vector.

(B) F nonnull: The Einstein-Maxwell questions for Eqs. (2) imply that (M_+, g_+) is a two-dimensional anti-deSitter space-time⁹ of radius A^{-1} , and that (M_-, g_-) is a 2-sphere of radius A^{-1} . This solution is the "Bertotti-Robinson magnetic universe"; for discussions see Bertotti,¹⁰ Robinson,¹¹ and Lindquist,¹² and Exercise 32.1 of Misner, Thorne, and Wheeler.¹³

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