

## Calculation of beta-decay half-lives of proton-rich nuclei of intermediate mass

G. T. Biehle and P. Vogel

*Physics Department, 130-33, California Institute of Technology, Pasadena, California 91125*

(Received 26 June 1992)

We present the results of a calculation of the beta-decay half-lives of several proton-rich even-even nuclei of intermediate mass:  $^{74}\text{Sr}$ ,  $^{76}\text{Sr}$ ,  $^{78}\text{Zr}$ ,  $^{80}\text{Zr}$ ,  $^{84}\text{Mo}$ ,  $^{86}\text{Mo}$ ,  $^{88}\text{Ru}$ ,  $^{90}\text{Ru}$ ,  $^{92}\text{Pd}$ , and  $^{96}\text{Cd}$ . The calculation is based upon the random phase approximation with the quasiparticle formalism and takes into account the residual particle-particle interaction.

PACS number(s): 23.40.Hc, 27.50.+e, 27.60.+j

Ever since Takahashi *et al.* [1] calculated in 1973 estimates for the beta-decay half-lives of virtually all beta-unstable nuclei, there has been a large effort to improve these estimates for neutron-rich nuclei because of applications in *r*-process theory and in the fate of fission products [2]. The theoretical effort to improve the estimates of half-lives of proton-rich nuclei has not been commensurate, although Hirsch *et al.* [3] and Muto *et al.* [4] presented calculations for light nuclei ( $Z \leq 30$ ). (See, however, also Refs. [5–7].) In this Brief Report we present the results of an effort to improve the estimates for half-lives of several proton-rich nuclei of intermediate mass. These half-lives play a role in *rp*-process theory [8], that is, the process in which protons are quickly added onto C, N, O, and other “metals” with intervening fast positron decays resulting in heavy proton-rich nuclei. This process occurs in certain astrophysical contexts in which the temperature is greater than about  $10^8$  K. In particular, this process is predicted to occur in massive stars with degenerate neutron cores (if they exist) [9], and information about the longer-lived ( $> 1$  s) beta-unstable nuclei would allow one to predict the nuclear abundances on the surfaces of these stars [10]. For this reason we undertook the calculation of half-lives of some proton-rich even-even nuclei of intermediate mass.

We are interested in even-even nuclei which have  $0^+$  ground states, so that the calculation is relatively simple. The positron-decay half-life  $t_{1/2}$  is given by the following formula:

$$\frac{1}{t_{1/2}} = \sum_m \frac{B(\text{GT})_m g_A^2}{6160 \text{ s}} f(\Delta E_m, Z), \quad (1)$$

where  $m$  labels the accessible  $1^+$  states in the daughter nucleus,  $B(\text{GT})_m$  is the Gamow-Teller  $\beta^+$  strength (equivalent to  $|\langle m | \sigma \tau^+ | i \rangle|^2$  in this case),  $g_A$  is the axial-vector-current coupling constant (which we set to 1.25), and  $f(\Delta E_m, Z)$  is the Fermi function (including Coulomb and relativistic corrections), which describes the size of phase space.

We obtain energy levels of the daughter nucleus and evaluate  $B(\text{GT})$  using the random phase approximation based on the quasiparticle formalism (QRPA). (The generalization of the QRPA to charge-changing modes is due to Halbleib and Sorensen [11]. Particle-particle interac-

tions were first included in the QRPA by Cha [12].) The formalism is described in detail in Vogel and Zirnbauer [13] and in Engel *et al.* [14]. In these papers the authors use the  $\delta$  force as the residual interaction and describe the following four parameters:  $\alpha_0$ ,  $\alpha_1$  (the particle-hole interaction constants in the  $S=0$  and  $S=1$  channels, respectively),  $\alpha'_0$ , and  $\alpha'_1$  (the particle-particle interaction constants). Although these constants are theoretically related, the authors present an argument that they can be treated independently in this calculation. Using the values given in Ref. [14], we set  $g_{\text{pair}} = -270 \text{ MeV fm}^3$  when we solve the BCS equations, and we set  $\alpha_0 = -890 \text{ MeV fm}^3$  and  $\alpha_1 = -1010 \text{ MeV fm}^3$  for the RPA portion of our calculations. Because we are looking at positron decay of proton-rich nuclei, our results do not depend on  $\alpha'_0$  in the RPA calculations. Our results do, however, depend strongly on the value of  $\alpha'_1$ , so we must take care to choose it carefully.

We divide the nuclei into two categories, those with  $74 \leq A \leq 80$  and those with  $80 < A \leq 96$ . For the heavier nuclei in our study, we calibrated  $\alpha'_1$  using the known decay half-lives of  $^{88}\text{Mo}$ ,  $^{90}\text{Mo}$ ,  $^{92}\text{Ru}$ , and  $^{94}\text{Pd}$ . In order to calculate these half-lives, we identified the lowest-lying  $1^+$  state in the daughter nucleus with the ground state given by the QRPA calculation. (This determines the values of  $\Delta E_m$ , used in the phase space integrals.) Our calculation is for positron decay only, i.e., no electron capture. In three of the calibration nuclei positron decay dominates over electron capture; however, 75% of the decay of  $^{90}\text{Mo}$  is due to electron capture. In that case we, therefore, use the proper partial decay rate. In our calculation, almost all ( $\geq 90\%$ ) of the predicted decays occur into the lowest-lying  $1^+$  state. Figure 1 shows the  $\log_{10}$  of the ratio of calculated positron-decay half-life to experimental half-life versus  $\alpha'_1$ . From this figure we see that  $\alpha'_1$  may be anywhere within a window from  $-324$  to  $-333 \text{ MeV fm}^3$  and yield values of half-lives correct to within a factor of 3. A value of  $\alpha'_1 = -329 \text{ MeV fm}^3$  yields a least  $\chi_{\text{red}}^2$  equal to 0.22, where

$$\chi_{\text{red}}^2 = \left[ \frac{1}{3} \sum [\log_{10}(T_{\text{calc}}/T_{\text{exp}})]^2 \right]^{1/2}.$$

Thus we predict that our results in Table I are accurate to about a factor of  $10^{0.22} = 1.7$ . By comparison, the  $\chi_{\text{red}}^2$

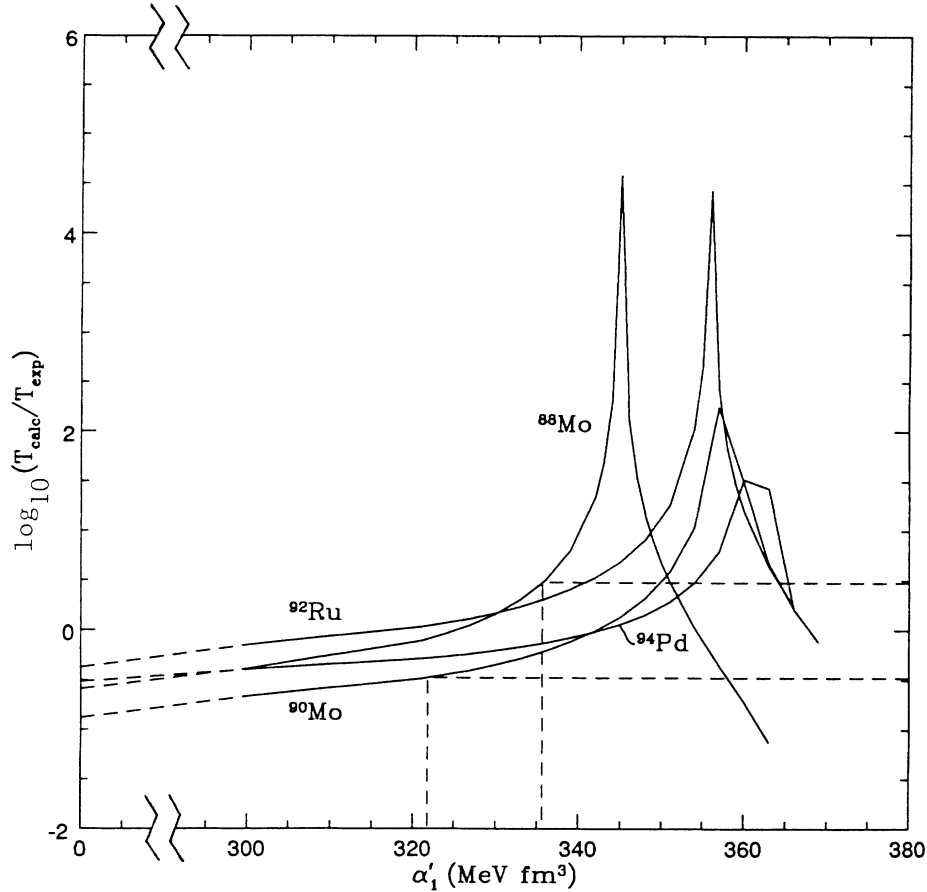


FIG. 1.  $\log_{10}(T_{\text{calc}}/T_{\text{exp}})$ , where  $T$  refers to the positron-decay half-life, versus  $\alpha'_1$  the particle-particle interaction strength. The window of values for  $\alpha'_1$  which yield results correct to within a factor of 3 is shown. Note that ignoring the residual particle-particle interaction (i.e., setting  $\alpha'_1$  to 0) results in prediction of half-lives approximately 3 to 10 times too small.

for these four nuclei using results from Takahashi *et al.* [1] is 0.59, yielding an estimated accuracy of a factor of  $10^{0.59} = 4$ .

In order to calculate half-lives of the nuclei listed in Table I, we need to know the positron-decay energies.

TABLE I. Predicted beta-decay half-lives.

| Nucleus          | $\Delta E_{m=0}^a$<br>(MeV) | Half-life <sup>b</sup><br>(s) | Takahashi <i>et al.</i> [1]<br>half-life (s) |
|------------------|-----------------------------|-------------------------------|--|
| <sup>74</sup> Sr | 9.6                         | 0.5                           | 0.03   |
| <sup>76</sup> Sr | 4.5                         | 8.                            | 3.   |
| <sup>78</sup> Zr | 10.5                        | 0.06                          | 0.03   |
| <sup>80</sup> Zr | 5.0                         | 7.                            | 3.   |
| <sup>84</sup> Mo | 5.2                         | 6.                            | 0.8  |
| <sup>86</sup> Mo | 3.9                         | 90.                           | 16.  |
| <sup>88</sup> Ru | 5.8                         | 1.2                           | 0.8  |
| <sup>90</sup> Ru | 4.7                         | 16.                           | 5.   |
| <sup>92</sup> Pd | 6.8                         | 0.9                           | 0.4  |
| <sup>96</sup> Cd | 8.0                         | 0.6                           | 0.3  |

<sup>a</sup> This is the maximum total energy of the positron for a transition to the lowest  $1^+$  daughter state.

<sup>b</sup> The estimated accuracy is a factor of 2. See the explanation in the text.

Since the masses of the positron-decay parents (and often those of the daughters as well) are not known, we use the predicted masses of Jänecke and Masson [15]. (These seem to reproduce best the known masses of proton-rich nuclei.) We set  $\Delta E_{m=0}$ , that is, the maximum total energy of the positron, to the difference of parent and daughter masses less 0.2 MeV. The 0.2 MeV represents a typical value for the energy difference between the ground state and the lowest-lying  $1^+$  state of the daughter nucleus. (For these decays, however,  $\Delta E_{m=0}$  is large enough that the correction is trivial.) The results are shown in Table I. As stated in the previous paragraph, these values are accurate to within a factor of about 2. Electron capture is negligible in these nuclei, contributing less than 3% because of the large decay energies involved. (See Ref. [16].)

Similarly we use the known half-lives of <sup>70</sup>Se, <sup>72</sup>Kr, <sup>74</sup>Kr, and <sup>80</sup>Sr to calibrate  $\alpha'_1$  and calculate half-lives for several nuclei with  $A \leq 80$ . In this case we obtain  $\alpha'_1 = -327 \text{ MeV fm}^3$  for the best fit, yielding a least  $\chi^2_{\text{red}}$  equal to 0.32. The results are also shown in Table I. We estimate that the results are accurate to within a factor of about  $10^{0.32} = 2$ , and again electron capture is negligible.

Also shown in Table I are the predicted half-lives of

Takahashi *et al.* It is encouraging that our results are consistent with theirs, which are calculated by a different method; most of the difference is due to different  $Q$  values (i.e.,  $\Delta E_{m=0}$ ), especially in the case of  $^{74}\text{Sr}$ .

The authors wish to acknowledge support from NASA Grant. No. NAGW-2920 and U.S. Department of Energy Contract No. DE-FG03-88ER40397.

- 
- [1] K. Takahashi, M. Yamada, and T. Kondoh, *At. Data Nucl. Data Tables* **12**, 101 (1973).
  - [2] A. Staudt, E. Bender, K. Muto, and H. V. Klapdor-Kleingrothaus, *At. Data Nucl. Data Tables* **44**, 79 (1990).
  - [3] M. Hirsch, A. Staudt, K. Muto, and H. V. Klapdor-Kleingrothaus, *Nucl. Phys.* **A535**, 62 (1991).
  - [4] K. Muto, E. Bender, and T. Oda, *Phys. Rev. C* **43**, 1487 (1991).
  - [5] J. Suhonen, *Phys. Lett. B* **255**, 159 (1991).
  - [6] I. N. Borzov, E. L. Trykov, and S. A. Fayans, *Yad. Fiz.* **52**, 985 (1990) [*Sov. J. Nucl. Phys.* **52**, 627 (1990)].
  - [7] A. Staudt, M. Hirsch, K. Muto, and H. V. Klapdor-Kleingrothaus, *Phys. Rev. Lett.* **65**, 1543 (1990).
  - [8] R. K. Wallace and S. E. Woosley, *Astrophys. J. Suppl.* **45**, 389 (1981).
  - [9] G. T. Biehle, *Astrophys. J.* **380**, 167 (1991).
  - [10] G. T. Biehle (unpublished).
  - [11] J. A. Halbleib and R. A. Sorensen, *Nucl. Phys.* **A98**, 542 (1967).
  - [12] D. Cha, *Phys. Rev. C* **27**, 2269 (1983); Ph.D. thesis, Michigan State University, 1982.
  - [13] P. Vogel and M. R. Zirnbauer, *Phys. Rev. Lett.* **57**, 3148 (1986).
  - [14] J. Engel, P. Vogel, and M. R. Zirnbauer, *Phys. Rev. C* **37**, 731 (1988).
  - [15] J. Jänecke and P. J. Masson, *At. Data Nucl. Data Tables* **39**, 265 (1988); see also pp. 289ff.
  - [16] C. M. Lederer and V. S. Shirley, *Table of Isotopes*, 7th ed. (Wiley, New York, 1978); see Appendix 21.