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Erratum: "Unlocking CO Depletion in Protoplanetary Disks. I. The Warm Molecular Layer" (2018, ApJ, 856, 85)

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The model dust distribution reported in the published article is incorrect. The published article states that dust grains with the same height as the gas, referred to as small grains, have an MRN size distribution of $r_d = 0.005$–$1\ \mu$m, while the grains that are more settled than the gas, i.e., large grains, have a size distribution of $r_d = 0.005$–$1000\ \mu$m. The correct dust distribution is as follows: there are two populations of dust grains with the same scale height as the gas. Fifteen percent of these grains have an MRN size distribution of $r_d = 0.005$–$1\ \mu$m, the remaining 85% have a size distribution of $r_d = 0.005$–$1000\ \mu$m. The large-grain population, which is more settled than the gas, has a size distribution of $r_d = 10$–$1000\ \mu$m. The results and conclusions remain unchanged.

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